

Rapid Optical/X-ray timing of black hole binaries: correlated and non-linear variability



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Optical power of X-ray binaries

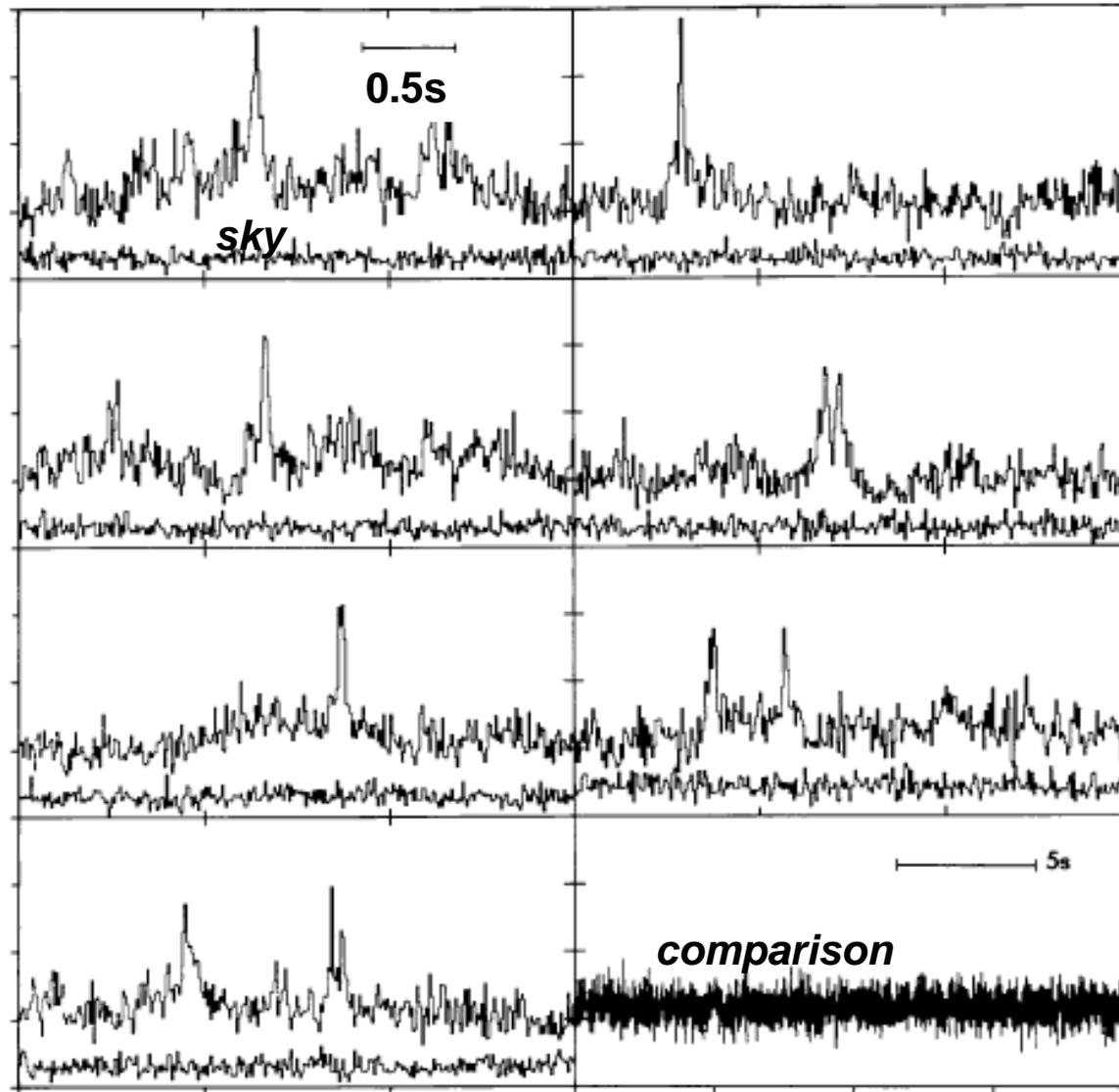


GX 339-4 ESO PR 2008

Artist: L. Calçada

Optical timing

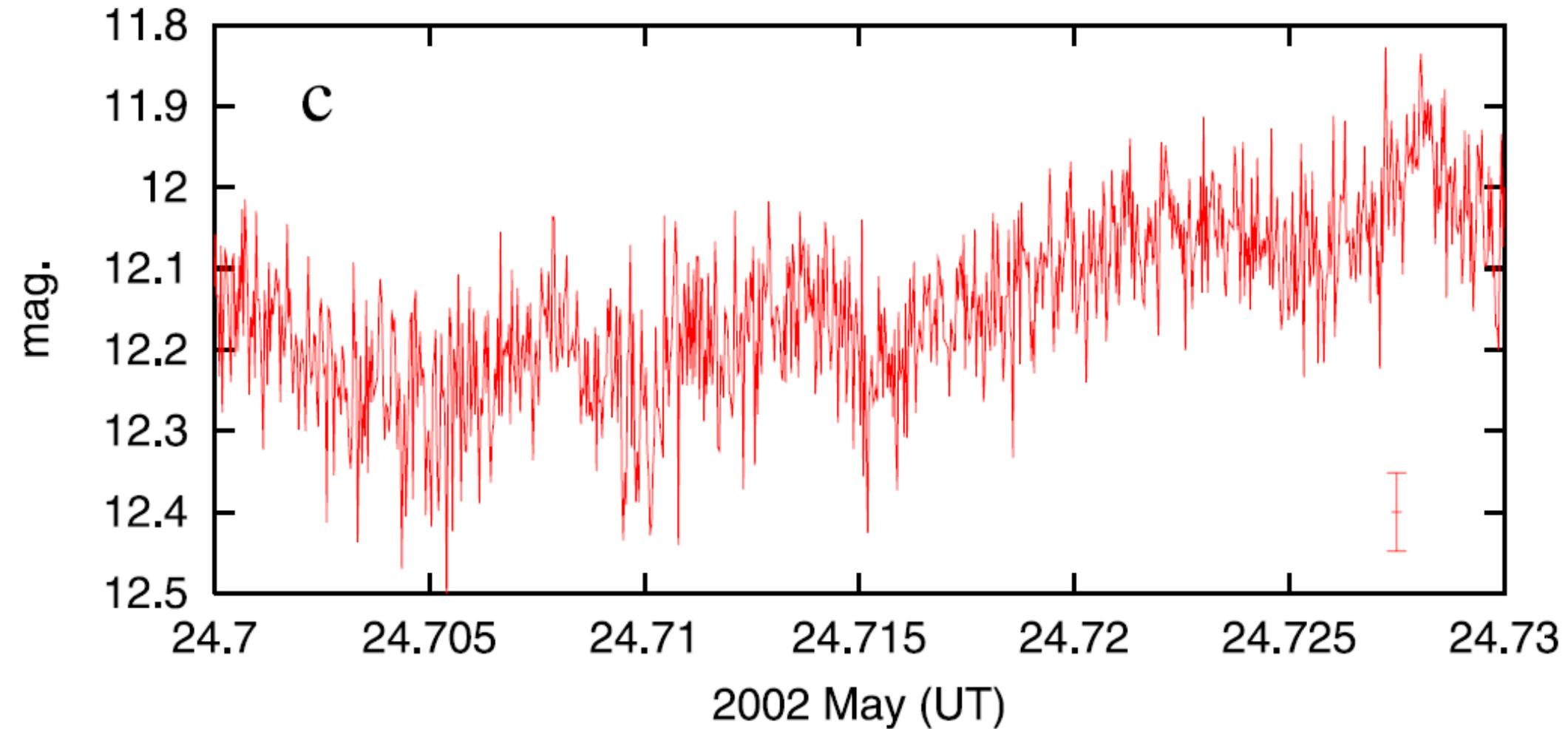
Early indications:
but not followed up, nor fully understood



(GX 339-4: Motch et al. 1982)

Optical timing

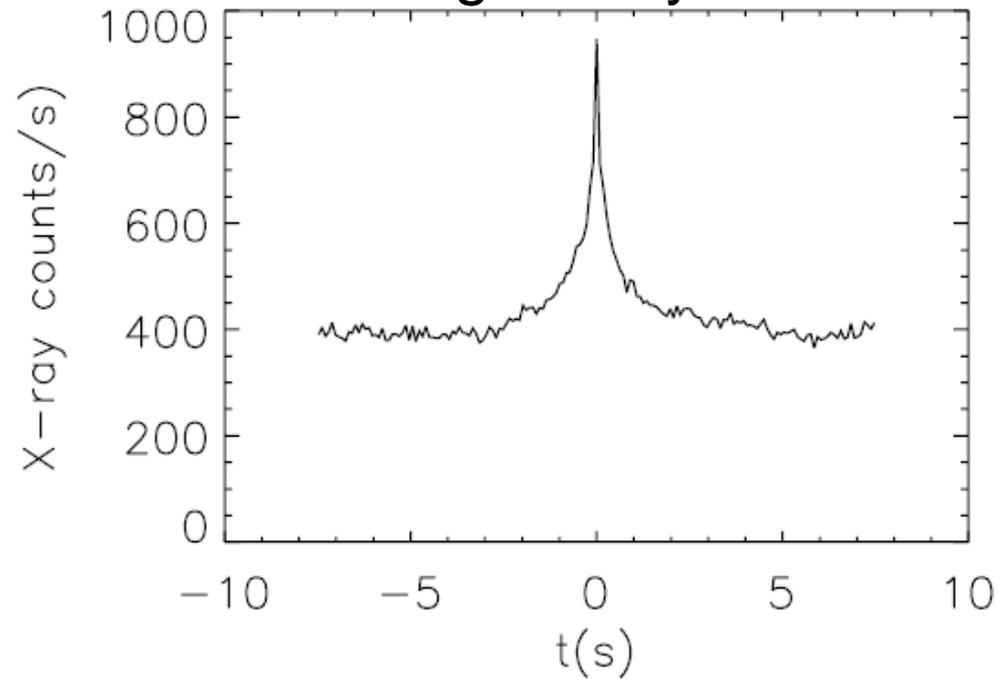
V4641 Sgr:



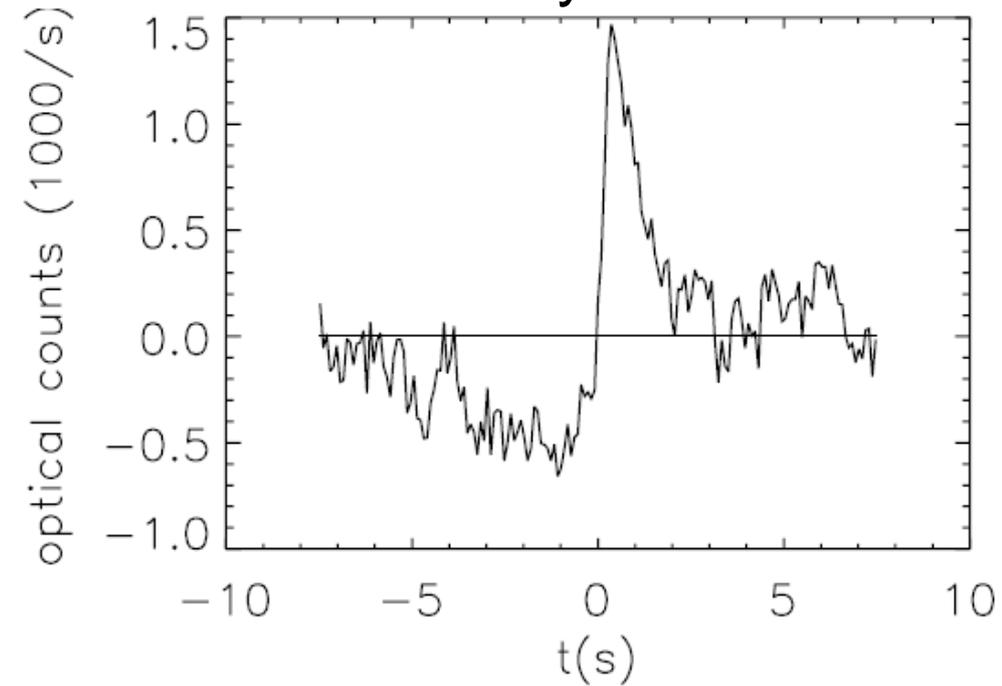
(Uemura et al. 2002)

XTE J1118+480

Average X-ray flares

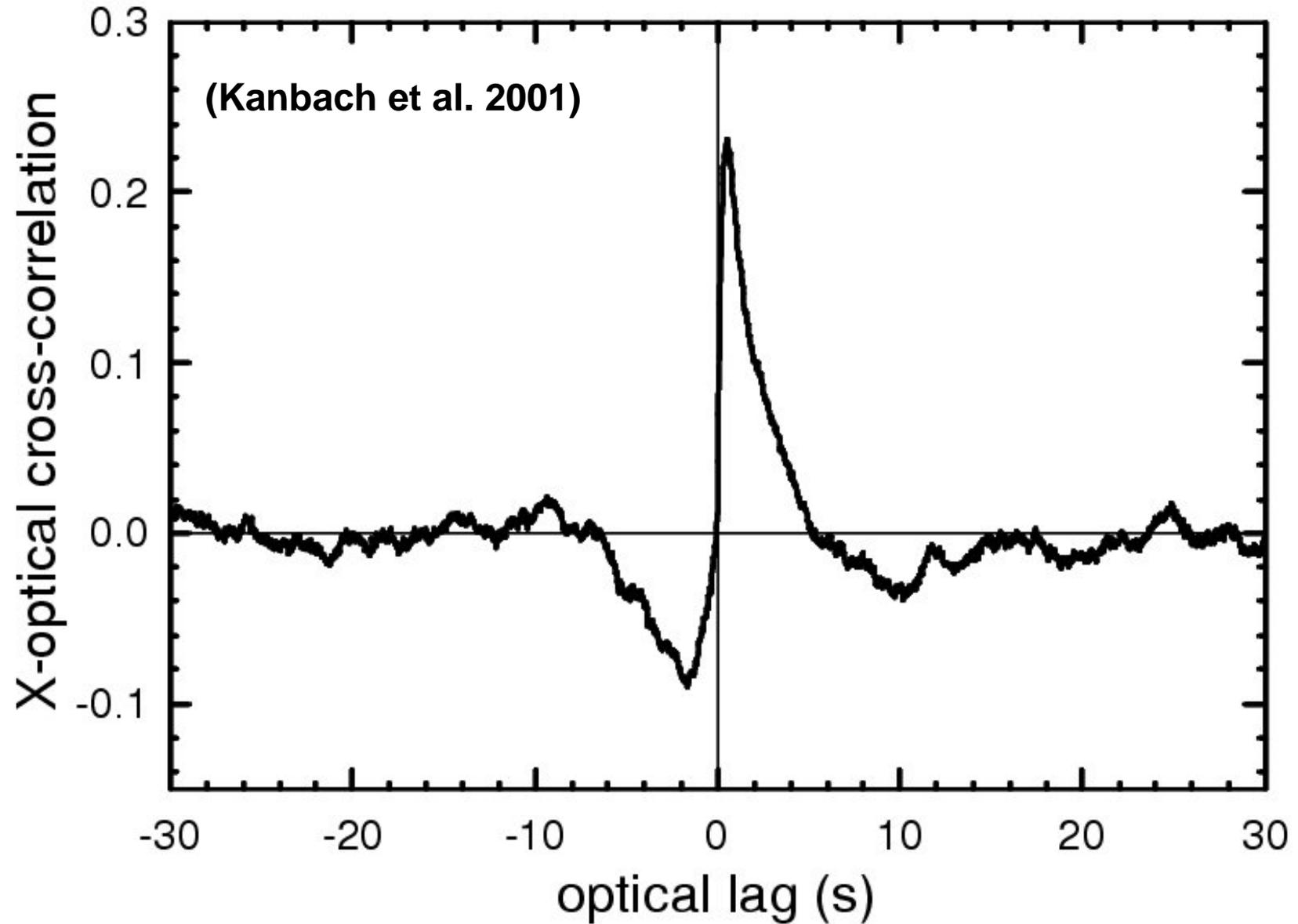


Average optical around X-ray flares



(Kanbach et al. 2001, Spruit et al. 2002)

XTE J1118+480



1996/1/8

2007/2/16

Our observations

Weighted mean flux (cts/s/SSC)



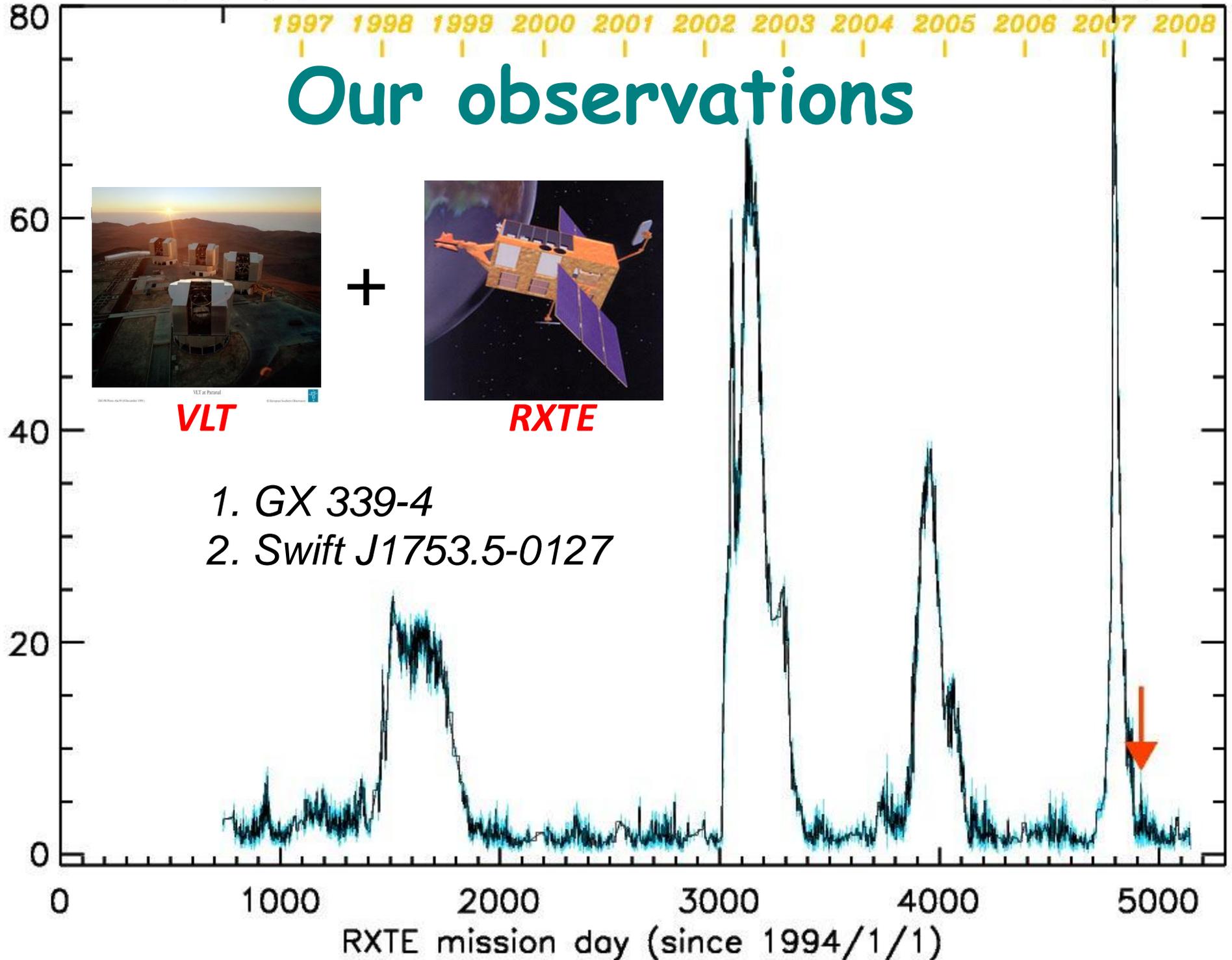
VLT

+



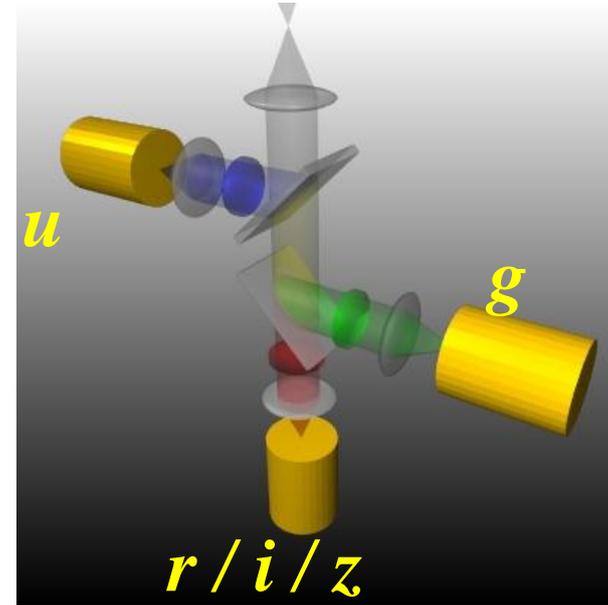
RXTE

1. *GX 339-4*
2. *Swift J1753.5-0127*



ULTRACAM: ultra-fast, triple-beam CCD camera

- Light-weight camera
(visitor instrument on *WHT/VLT/NTT*)
- Frame-transfer CCDs with negligible dark current, dead-time
- Speeds ~ 500 frames / sec
- 3 simultaneous optical filters
- Absolute timing ~ 1 ms



ULTRACAM Mounted on Visitor Focus of MELIPAL

ESO PR Photo 19a/05 (9 June 2005)

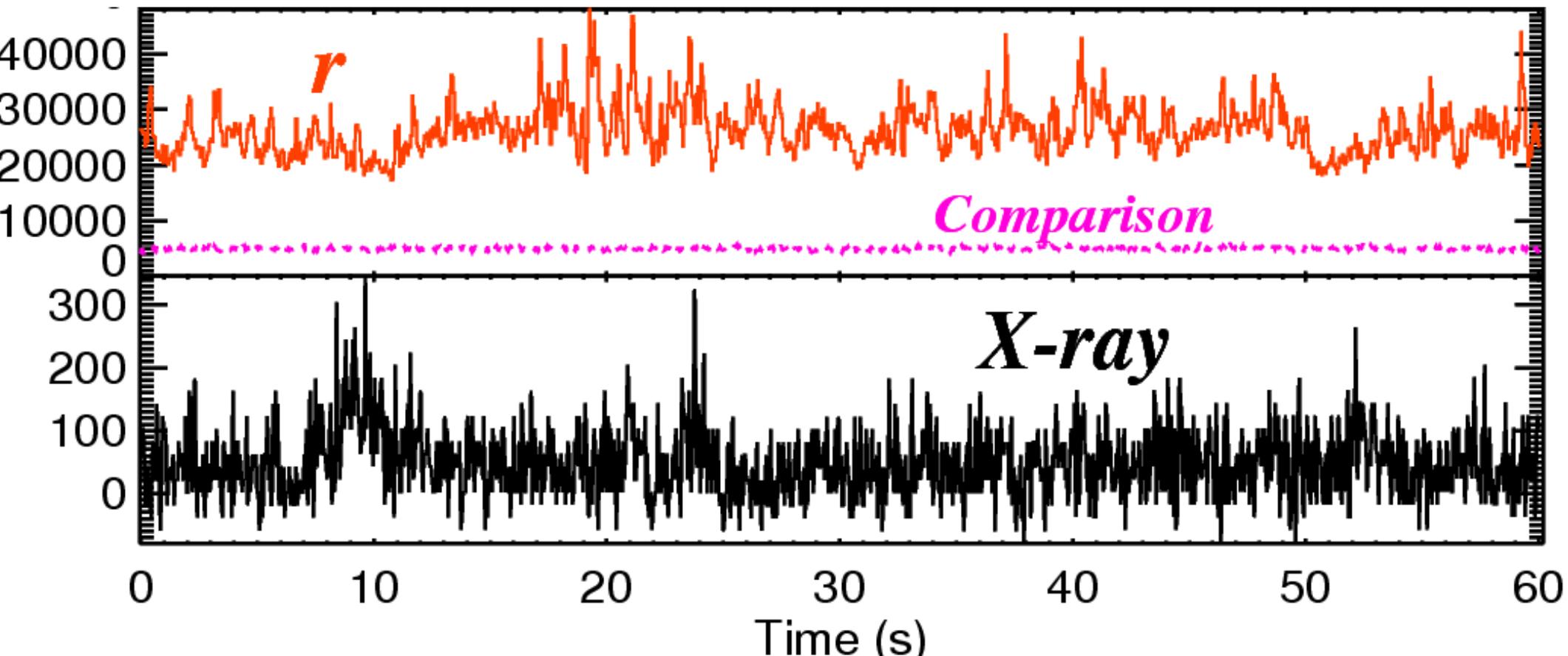
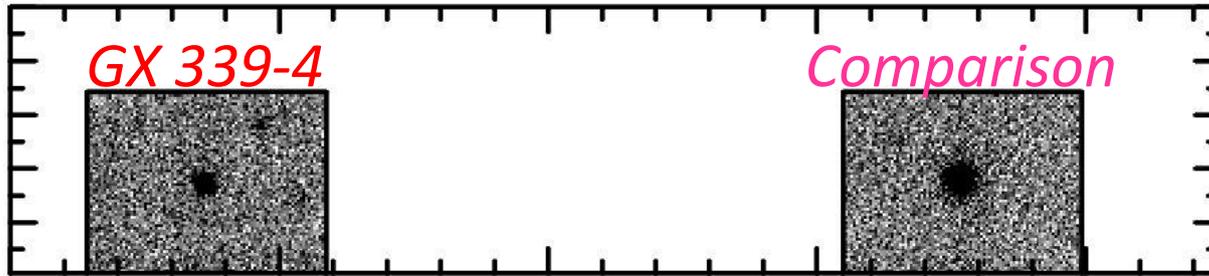
© ESO

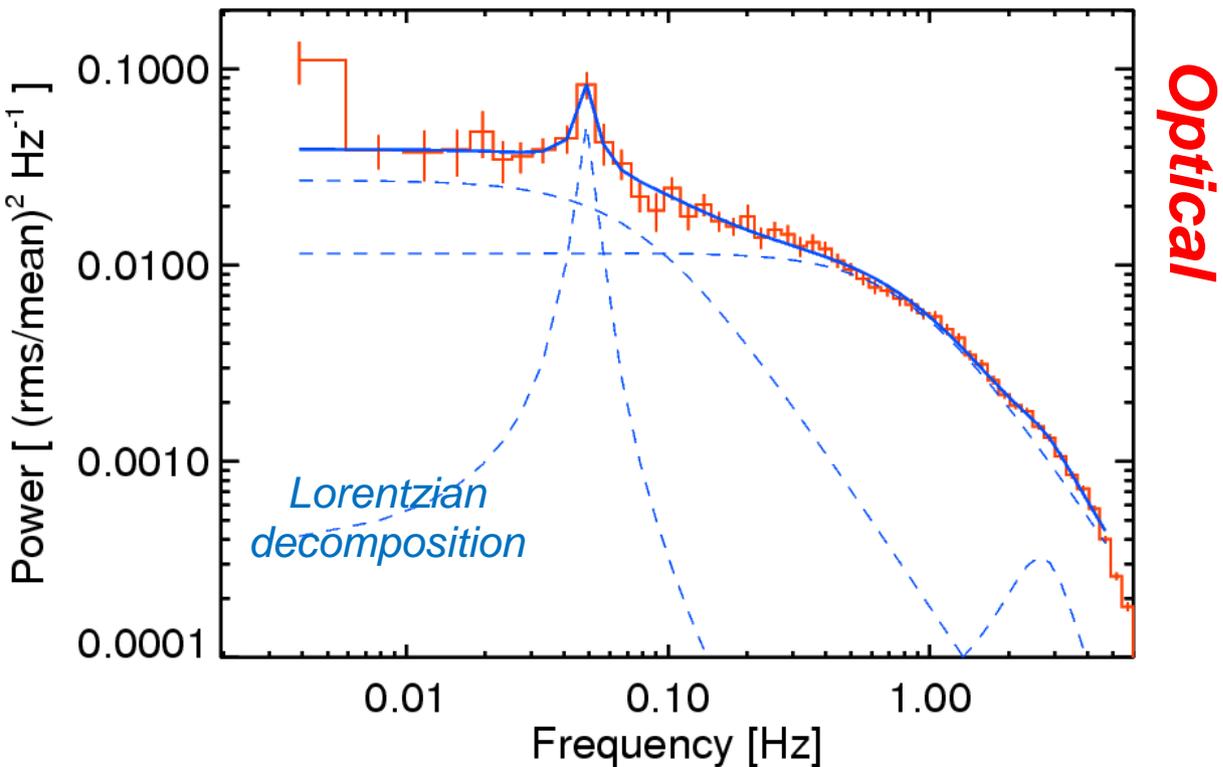


<http://www.shef.ac.uk/physics/people/vdhillon/ultracam/>

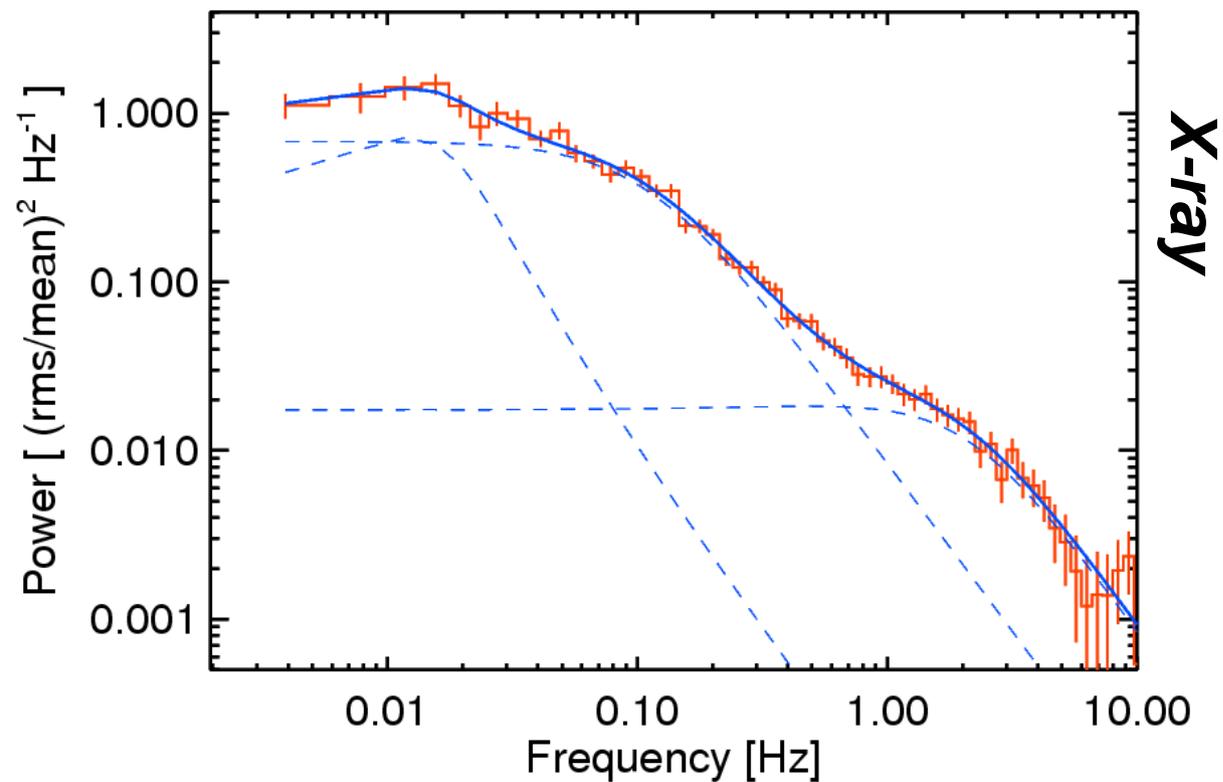
Observations

50 ms ULTRACAM frame (r' band)



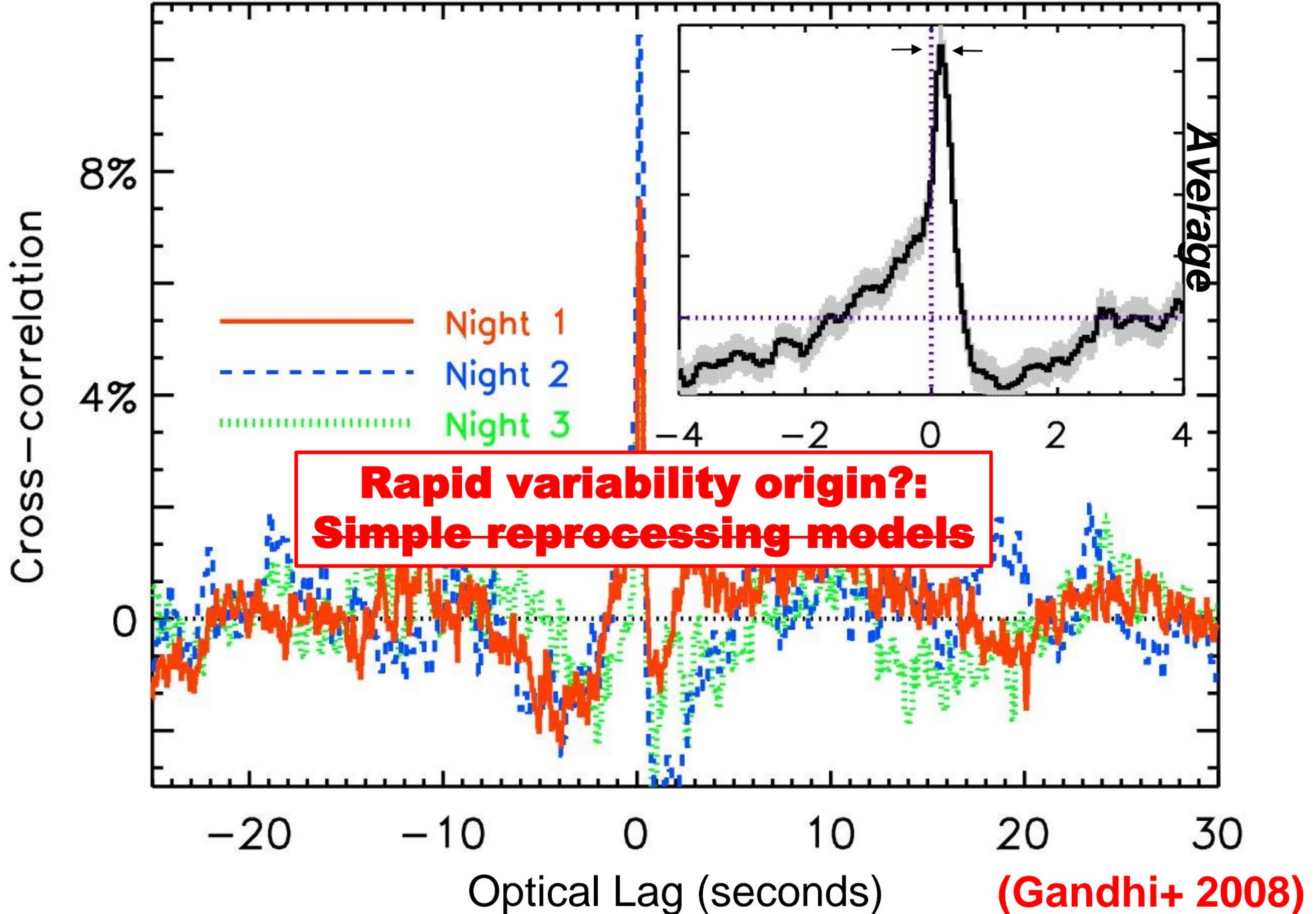


**Power
Spectra
show**

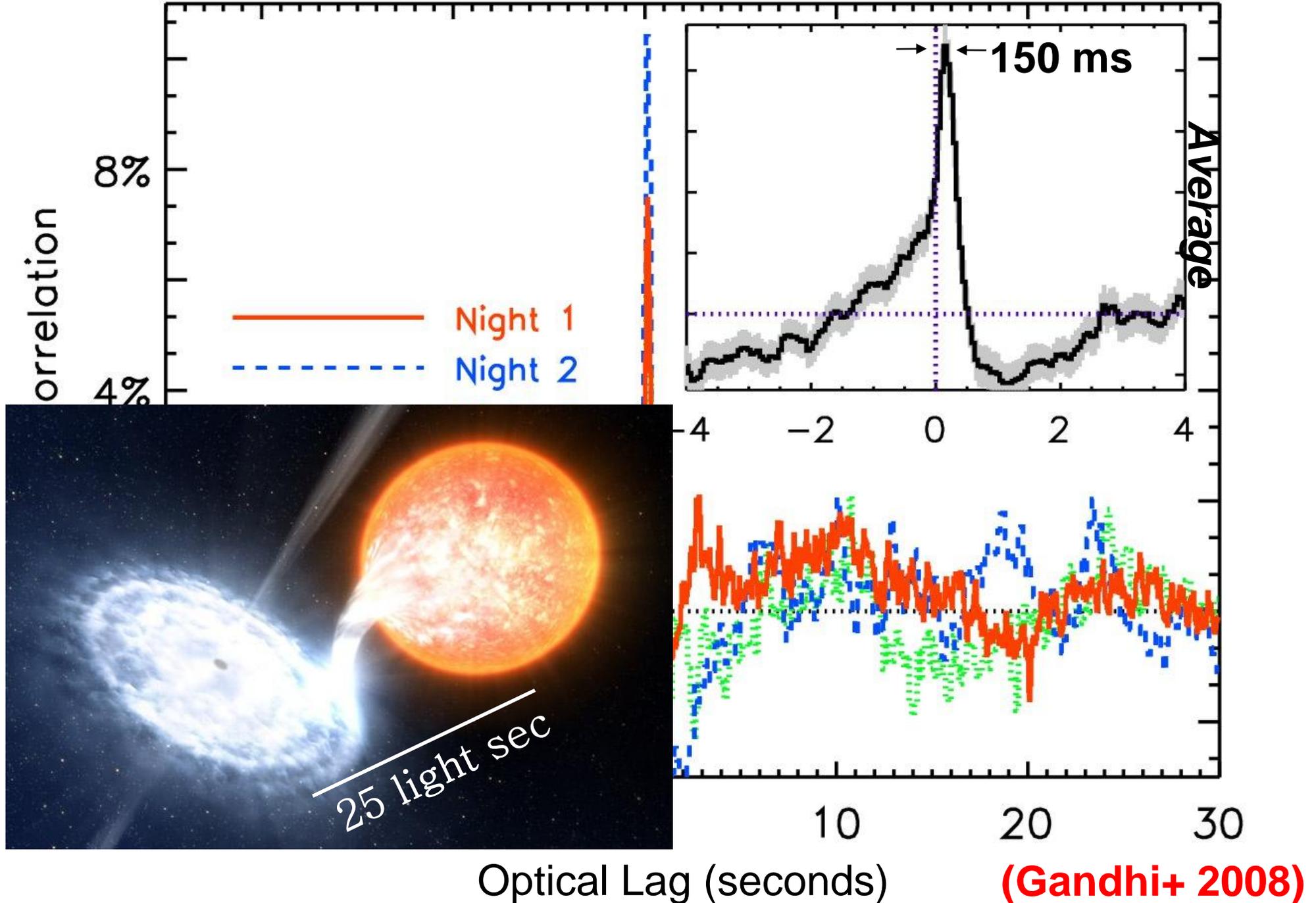


**broad-band
noise**

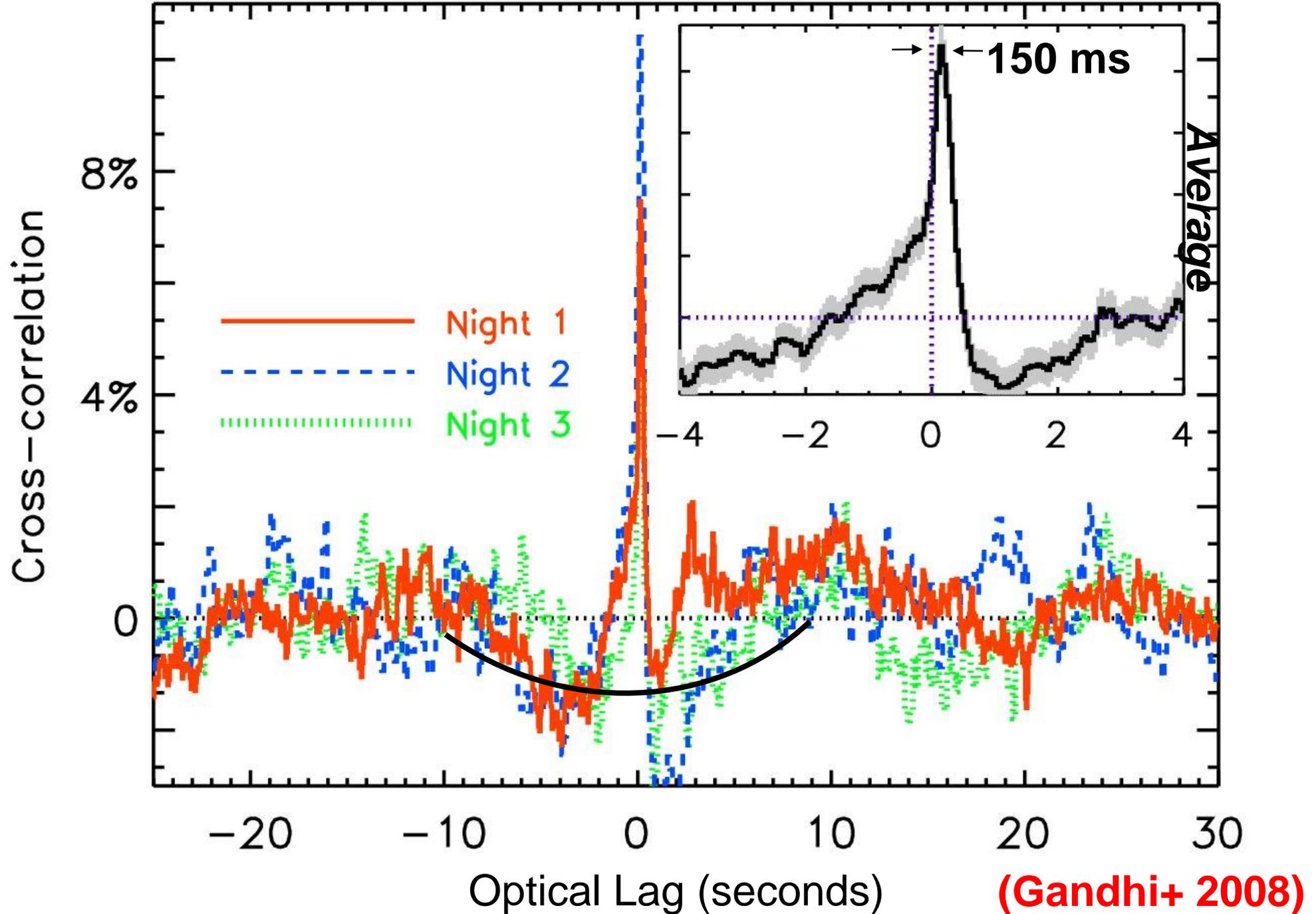
X-O Cross Correlation Function (CCF)



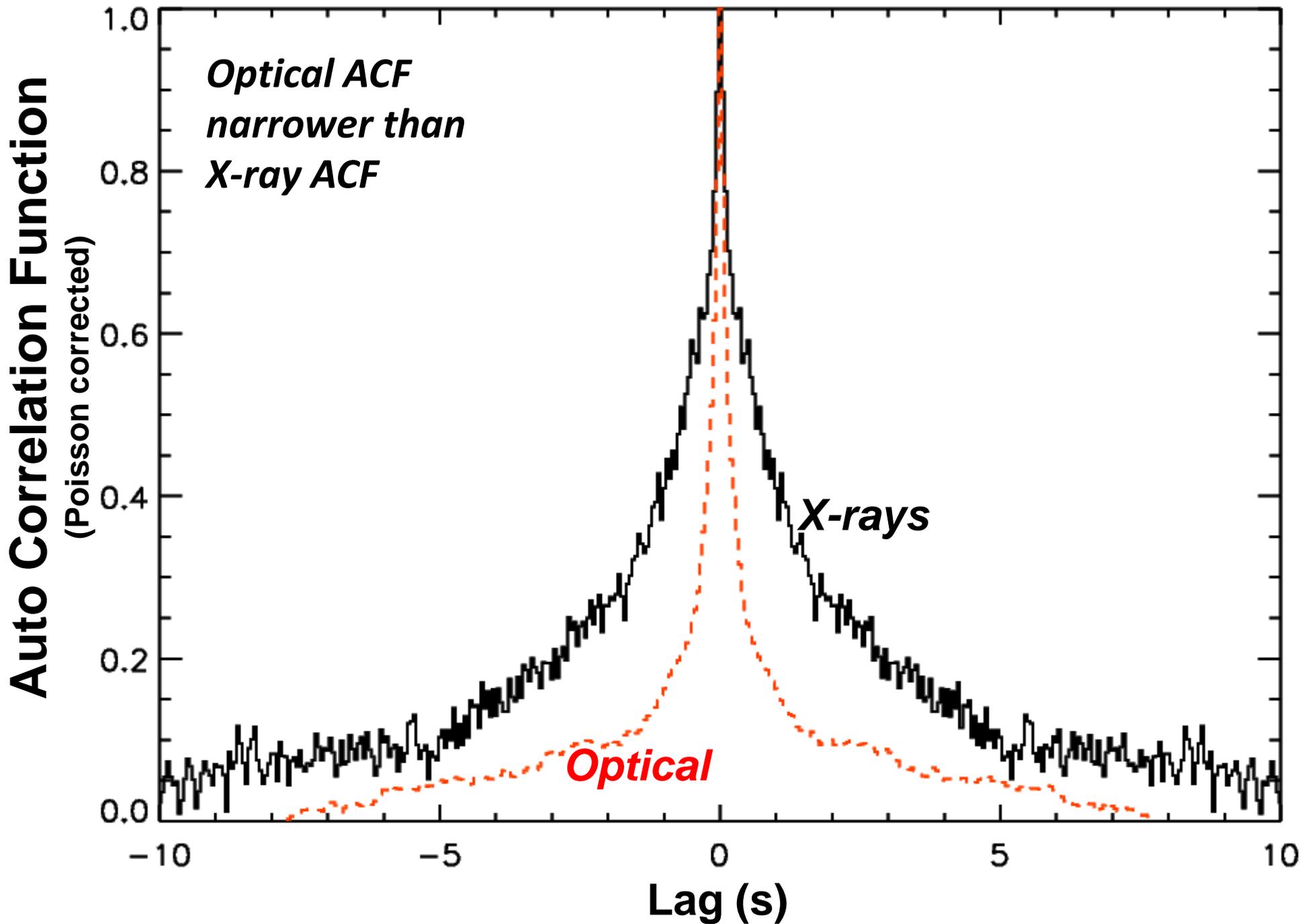
1. Small time delay



2. Anti-correlation

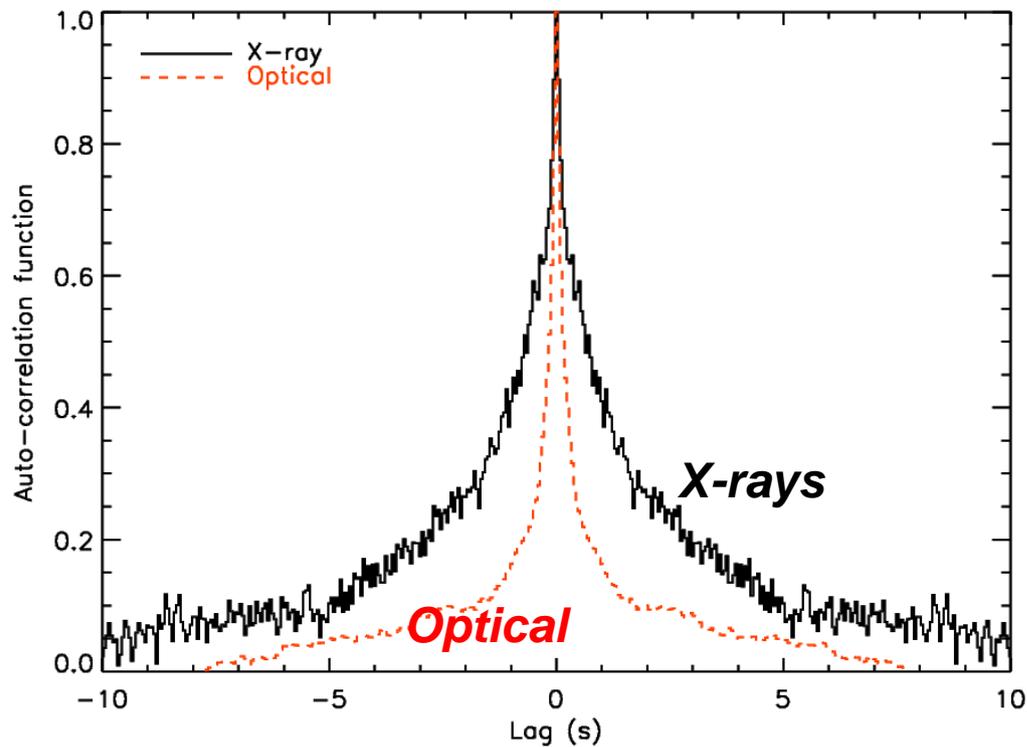


3. Small optical coherence times

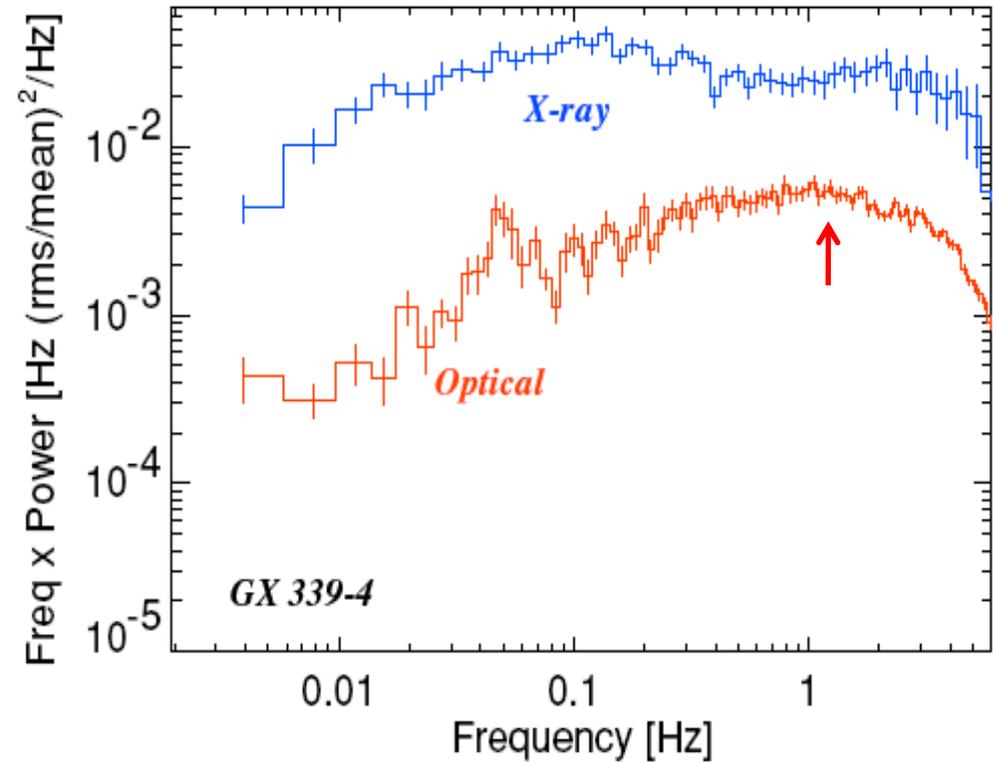


3. Small optical coherence times

Auto correlation functions

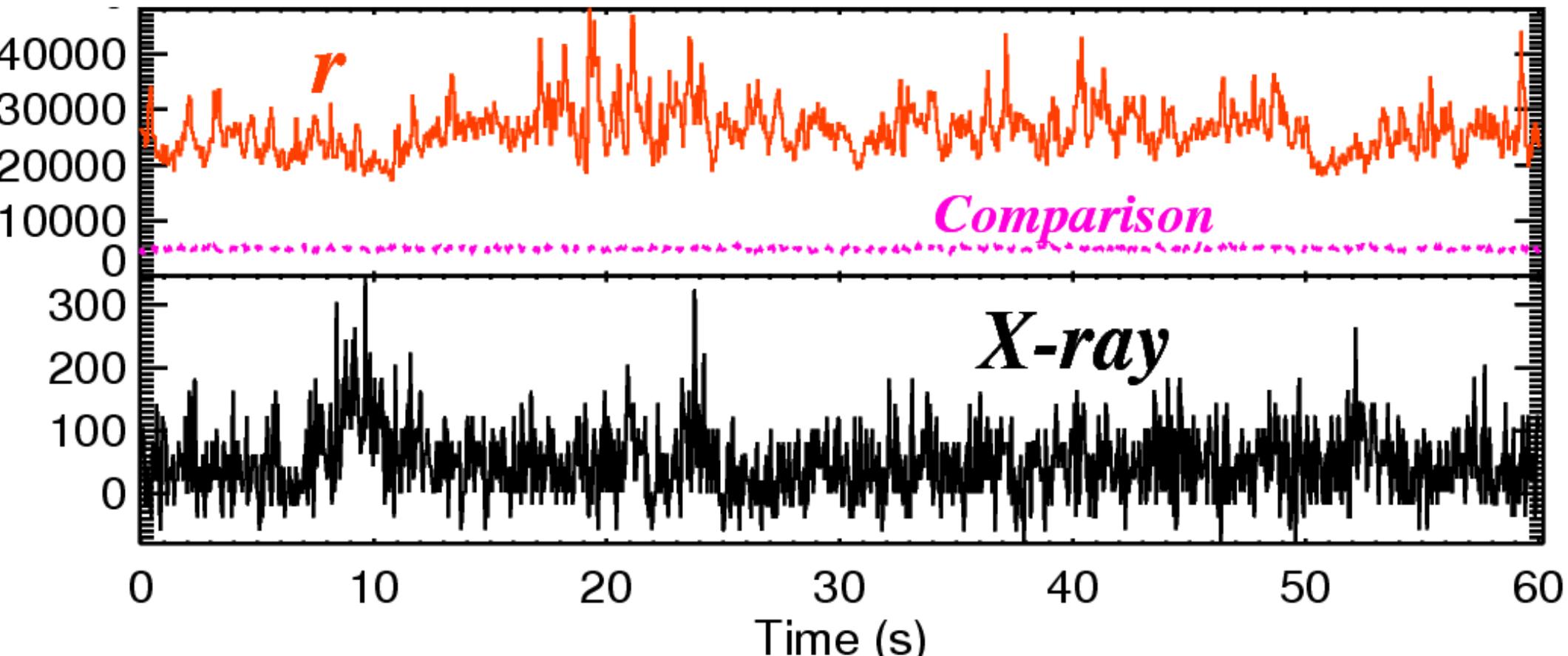


Power spectra

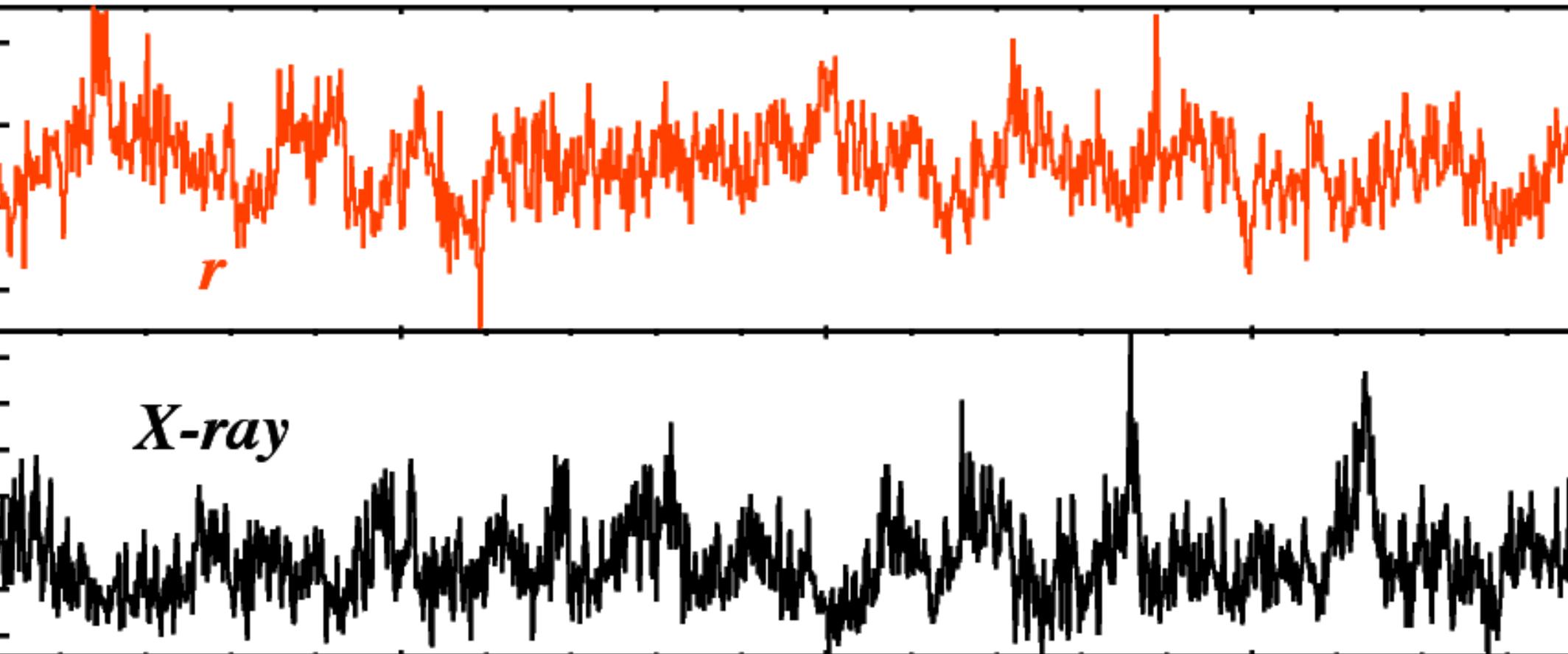


(rms² normalization: Belloni+90)

GX 339-4: Simultaneous light curves

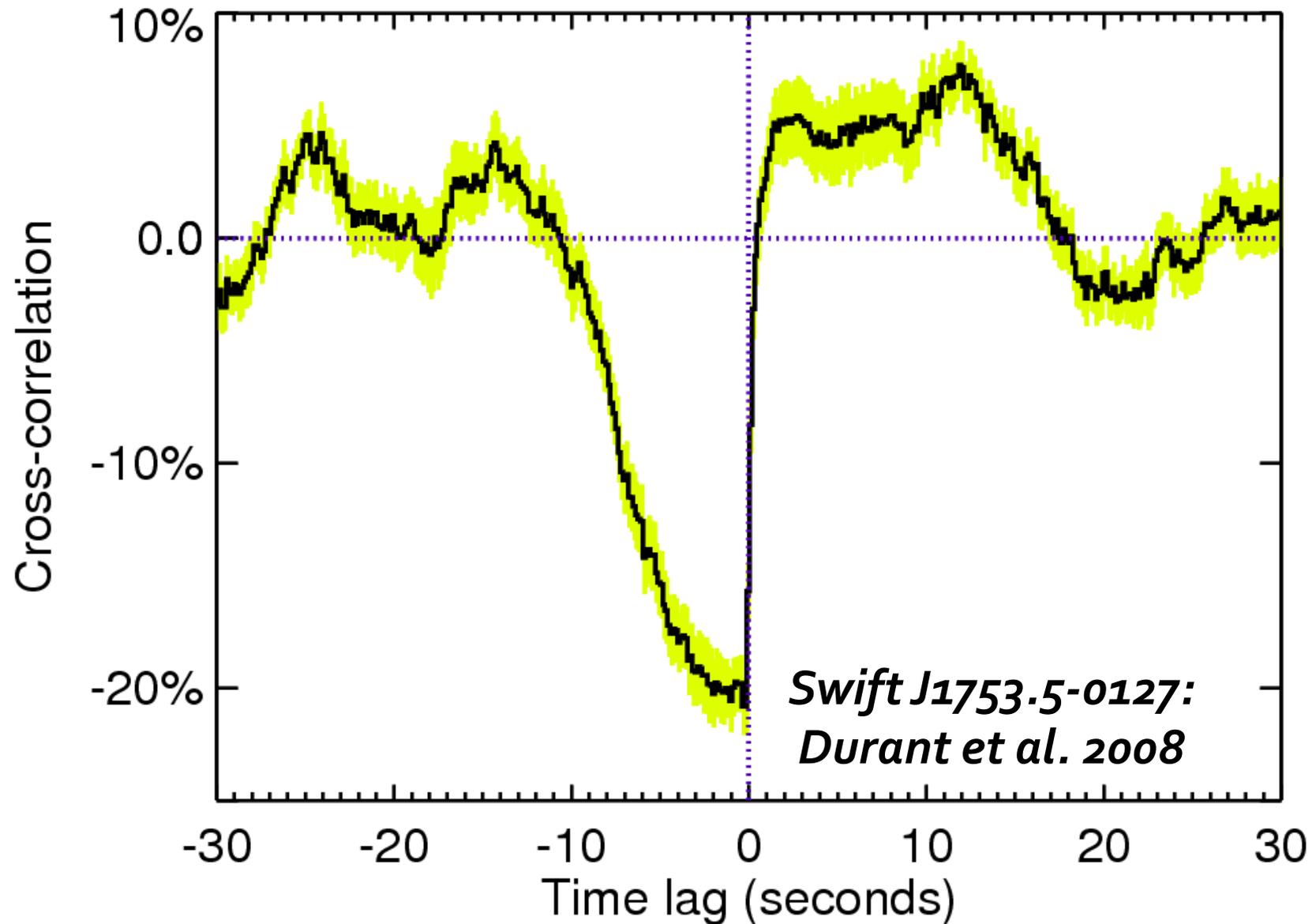


Swift J1753.5-0127: Simultaneous light curves



Time (s)

Complex optical/X-ray correlations

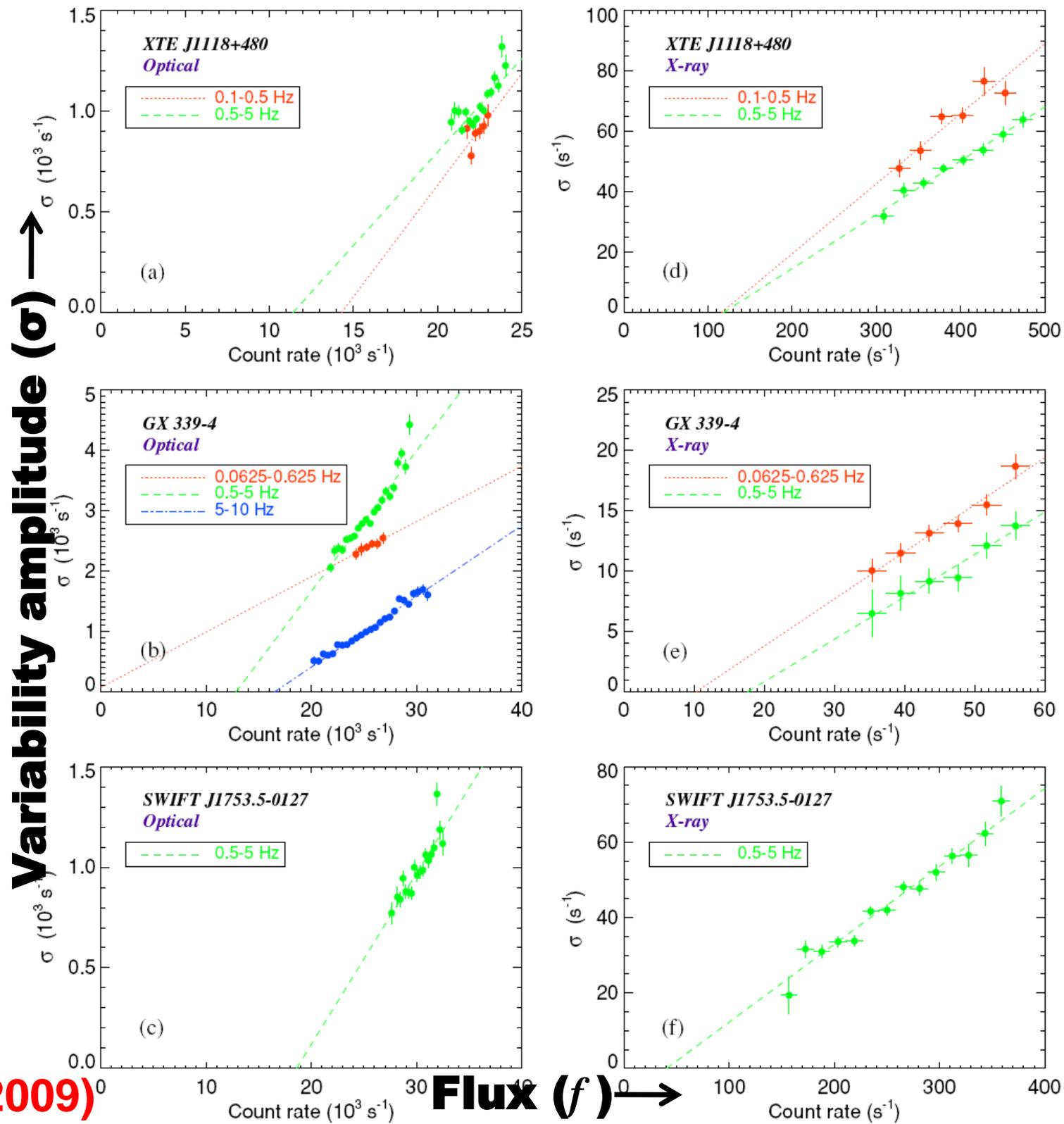


Non-linear variability:

$$\frac{df}{dt} \propto f$$

Random shot noise predicts constant σ

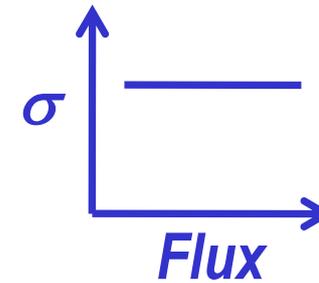
(Gandhi 2009)



Additive shots ruled out

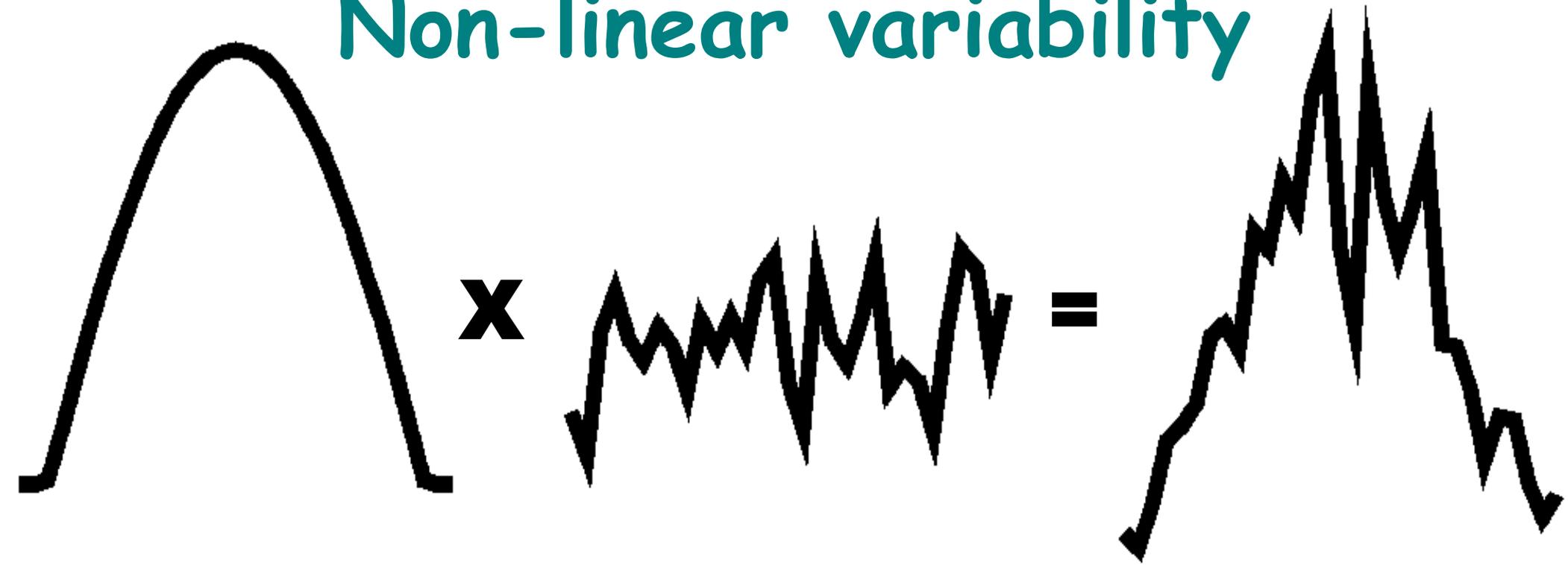


Superposition of independent shots \Rightarrow



**Ruled out in X-rays (Uttley et al. 2001...2005)
and now optical (Gandhi 2009)**

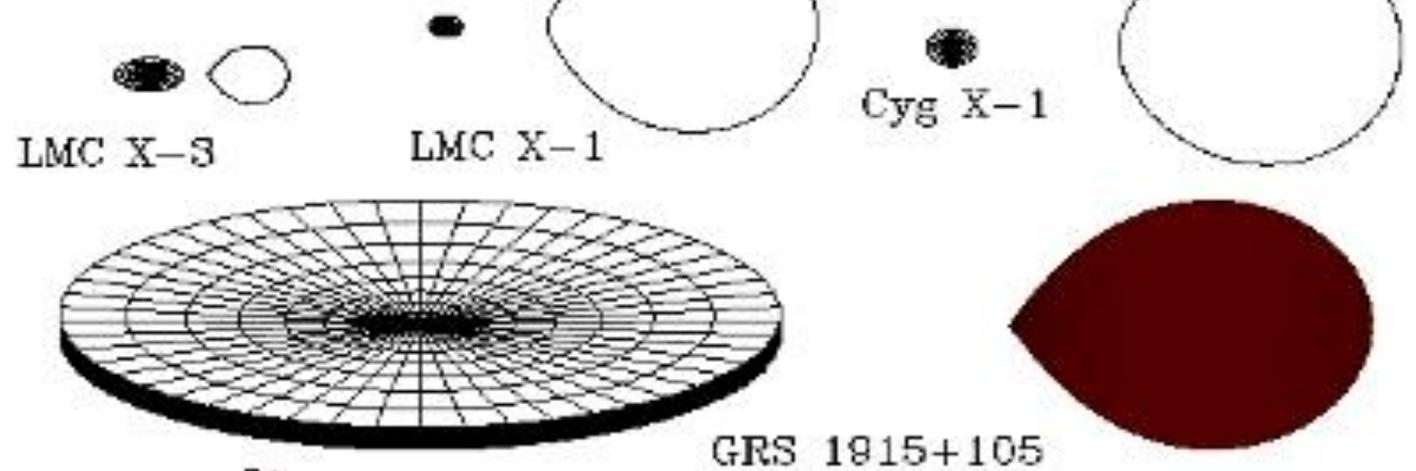
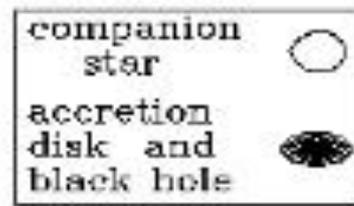
Non-linear variability



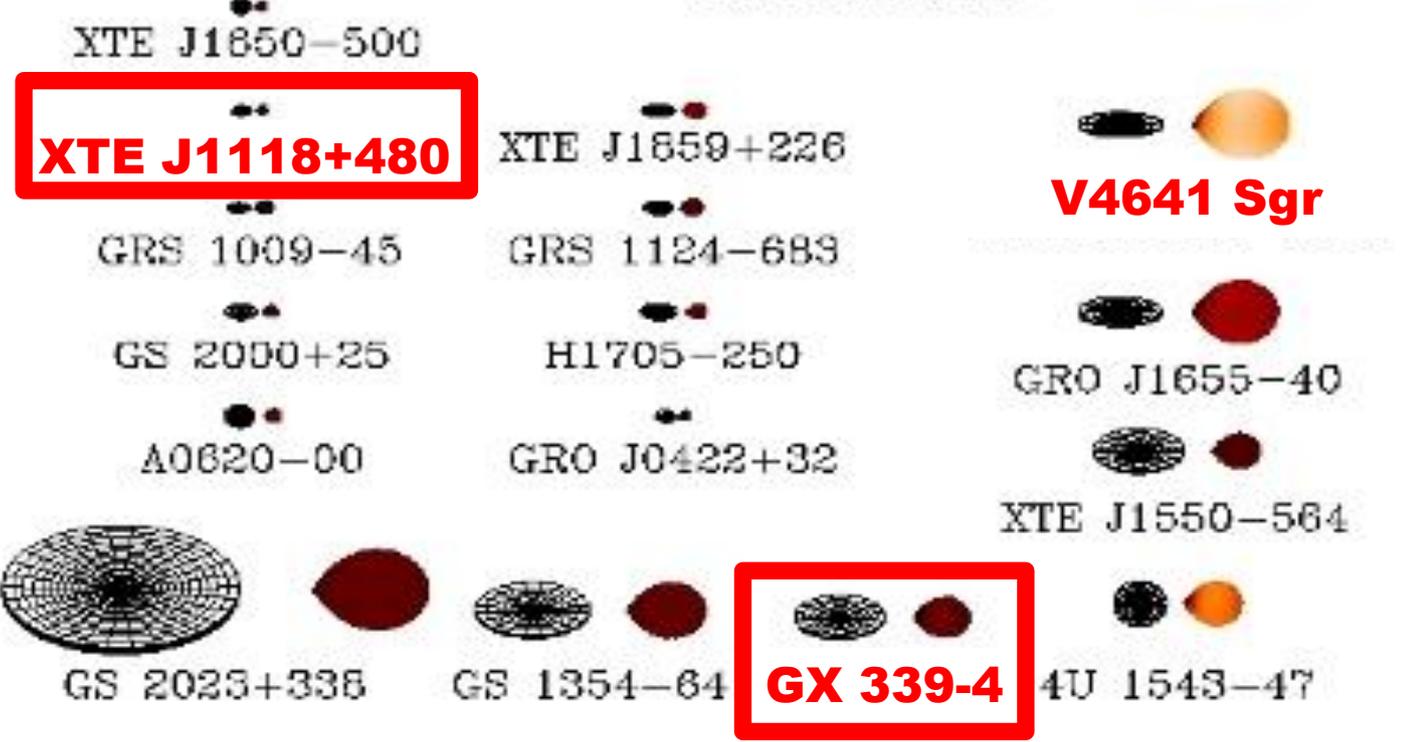
Perturbations must be *coupled* together, rather than *superposed*



Interactions between multiple emission components
(Lyubarskii 97, Misra 00, King+04, Titarchuk+07, Zhang 07)



X-ray binaries

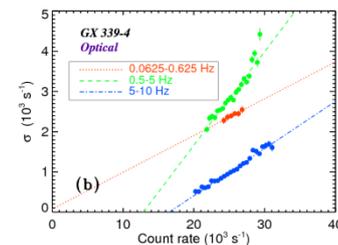
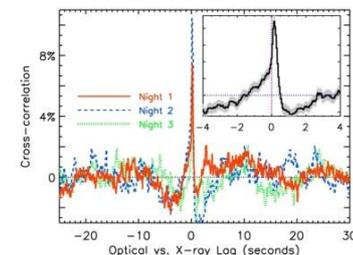
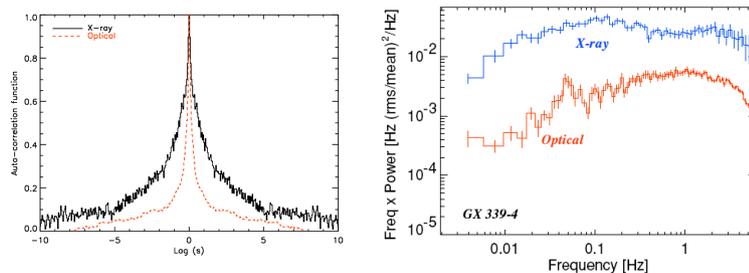
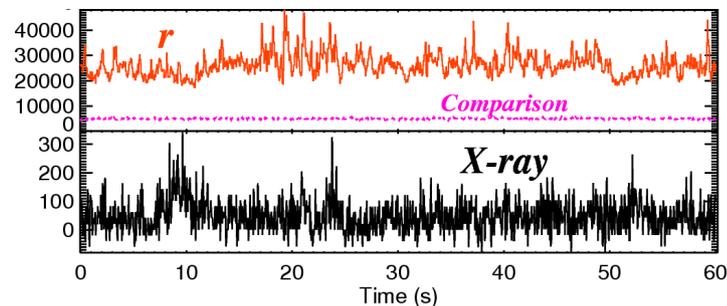


(J. Orosz)

Swift J1753.5-0127

Summary

- Simultaneous rapid optical/X-ray timing of X-ray binaries in low/hard state.
- Optical not reprocessed.
- Complex CCF
=> jet/corona/disk interaction
- X-ray and optical r.m.s. scales with flux
=> additive shots ruled out



Fast optical timing => interesting constraints
on accretion and hot plasmas