

Infrared Excess sources: Compton Thick QSOs at $z \sim 2$?

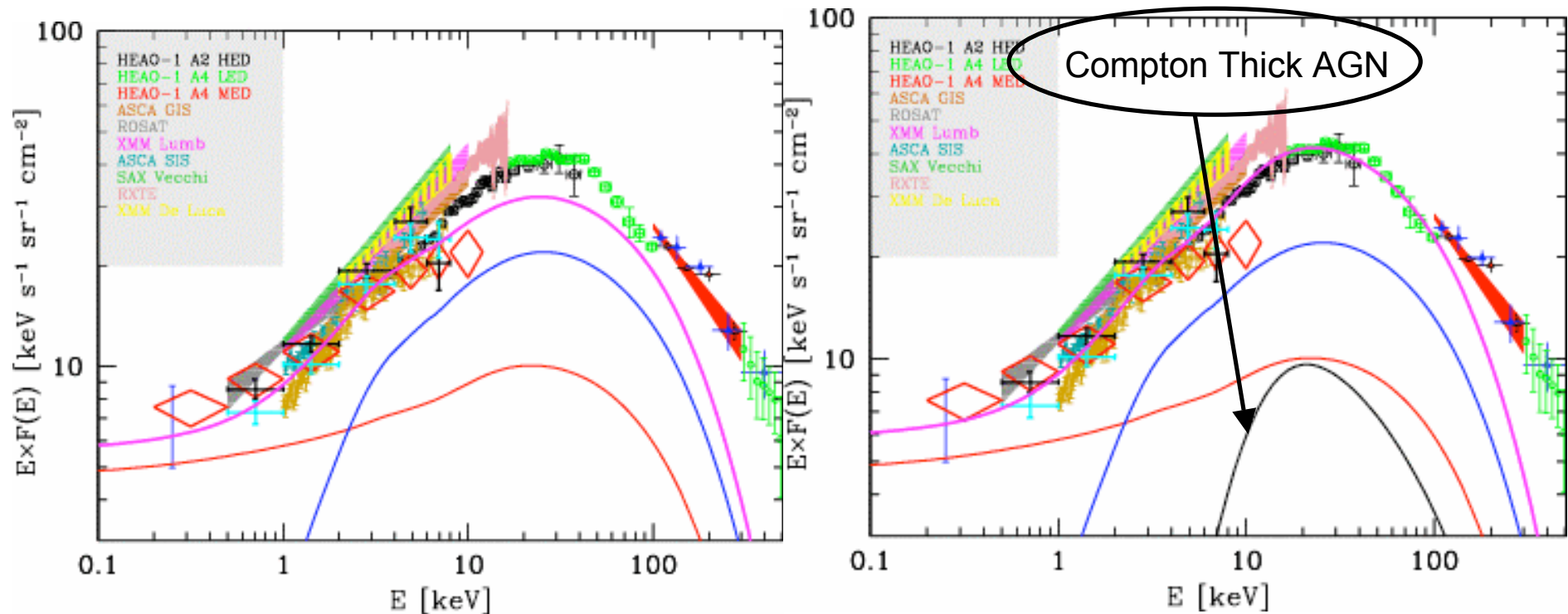


Antonis Georgakakis

National Observatory of Athens

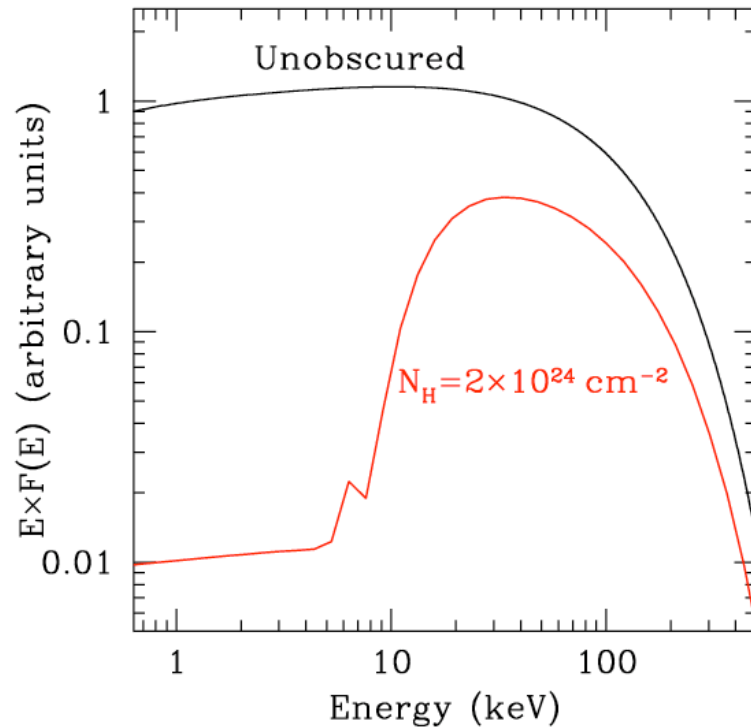
K. Nandra, M. Rowan-Robinson (Imperial College, London),
P. G. Pérez-González (Universidad Complutense de Madrid)

X-ray Background models

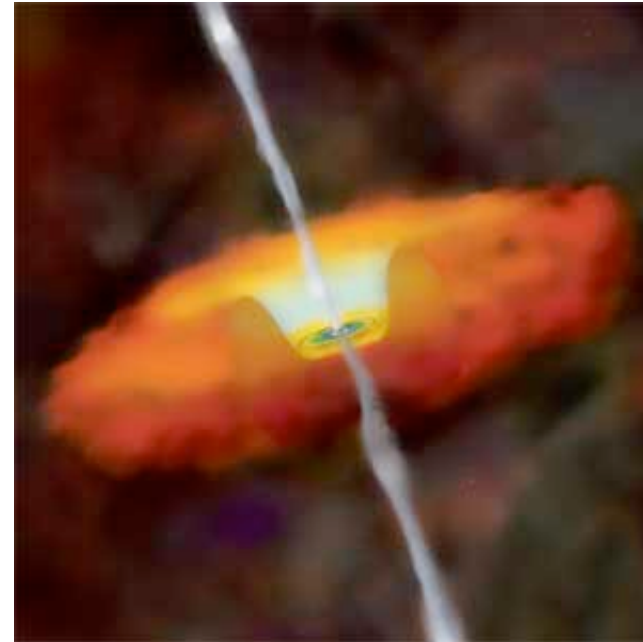


Compton Thick AGN ($N_H > 2 \times 10^{24} \text{cm}^{-2}$)
required by models for the XRB

Finding Compton Thick AGN: X-ray vs mid-IR

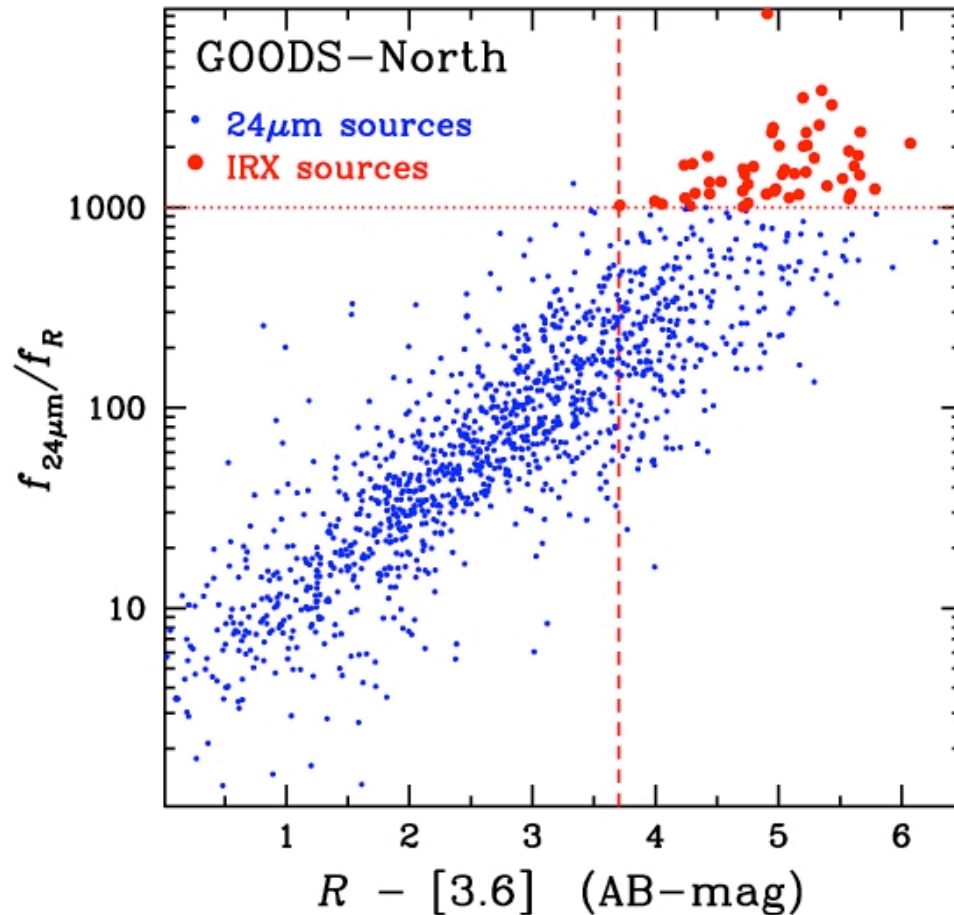


Hard to detect below
 $\sim 10 \text{ keV}$



mid-infrared ($3\text{--}20 \mu\text{m}$):
alternative wavelength
regime for finding
obscured AGN

Infra-Red EXcess source selection

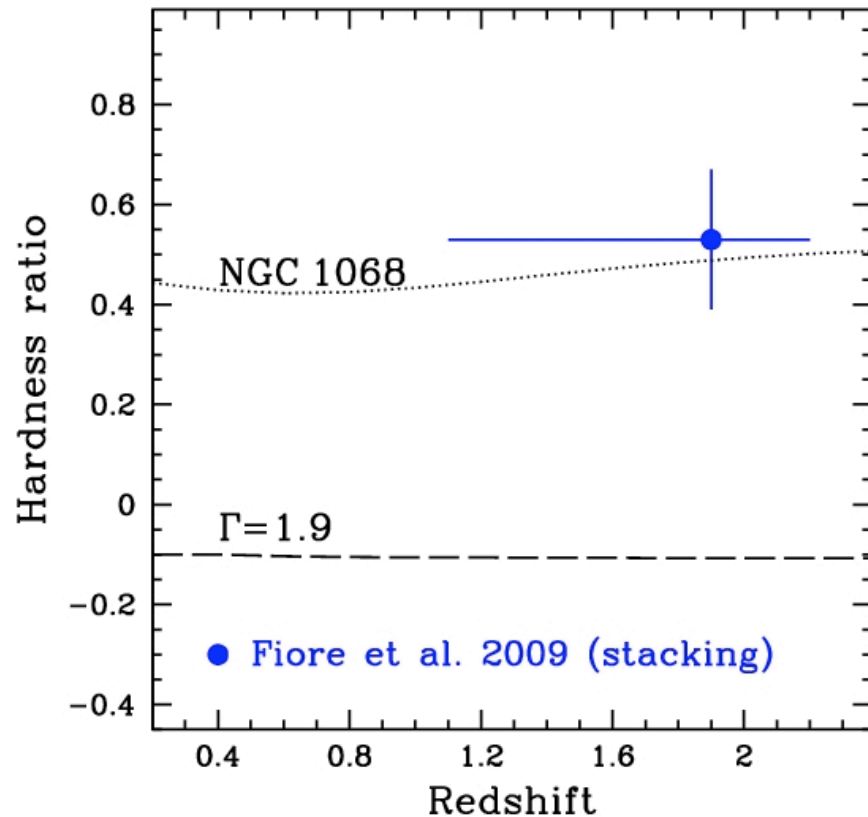


- $f_{24\mu\text{m}}/f_R > 1000$
- $R - [3.6] > 3.7$
- mean redshift, $z \sim 2$

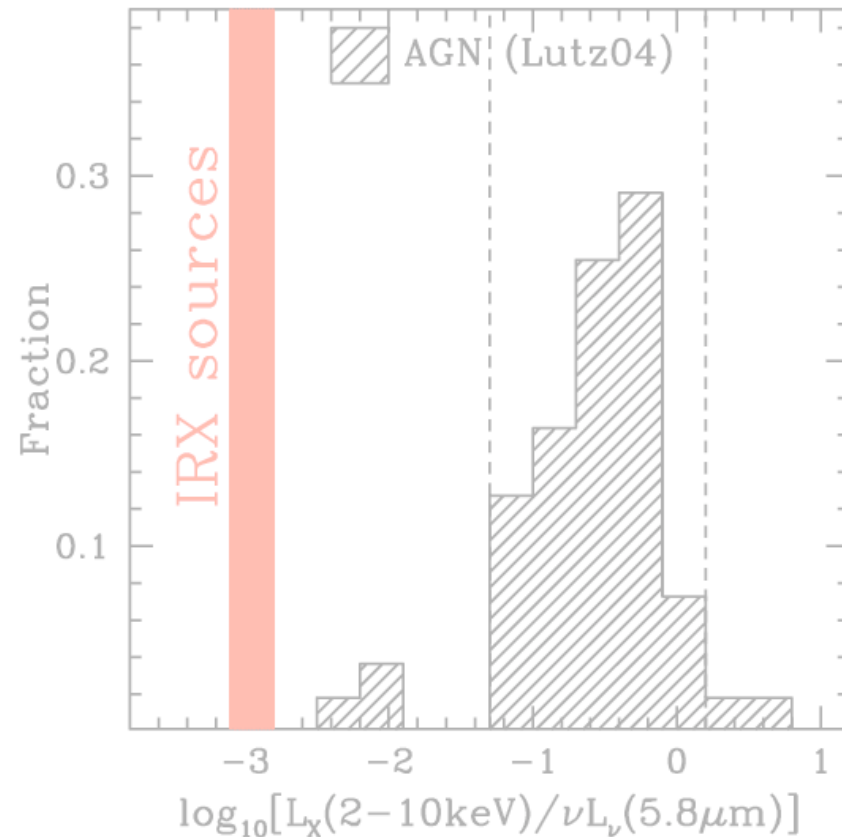
Daddi et al. 2007, Fiore et al.
2008, 2009 (IRX sources)

Dey et al. 2008 (DOGs)

IRX sources: Compton Thick QSOs?

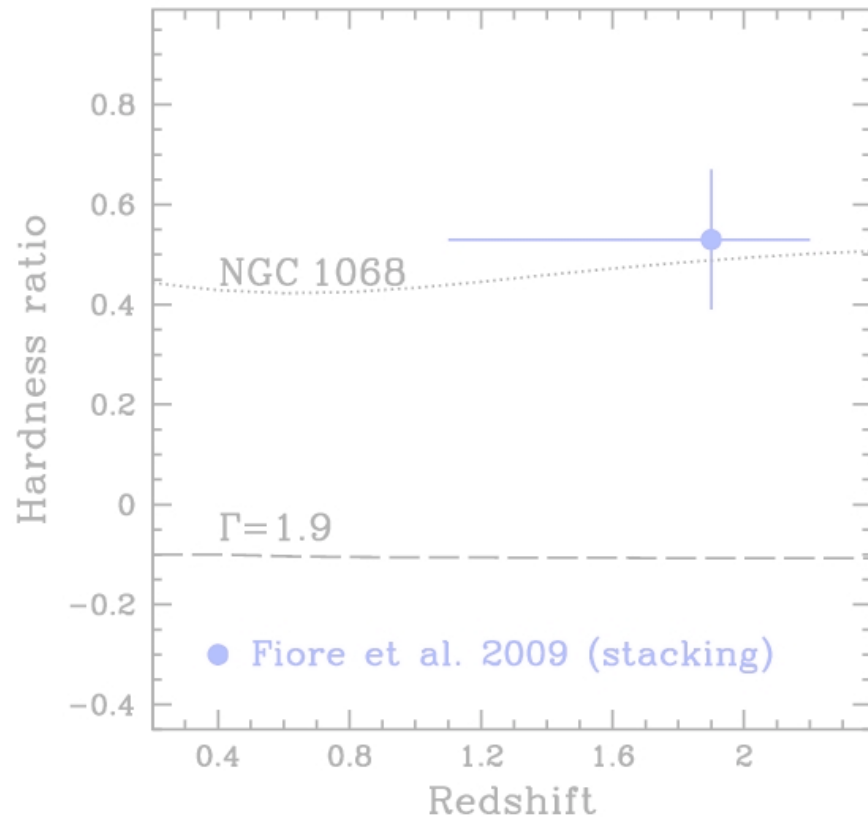


X-ray stacking: hardness ratio consistent with Compton Thick AGN ($N_H > 2 \times 10^{24} \text{ cm}^{-2}$).

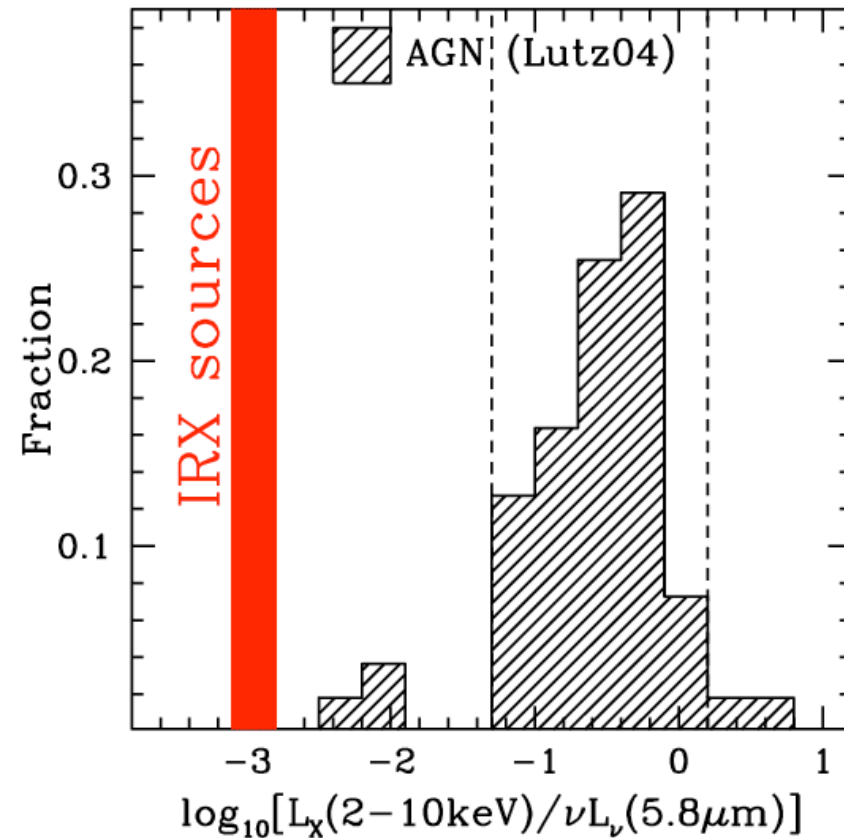


X-ray faint: 2-3dex fainter than typical AGN. Evidence for CT QSOs ($L_X \gg 10^{43} \text{ erg/s}$).

IRX sources: Compton Thick QSOs?

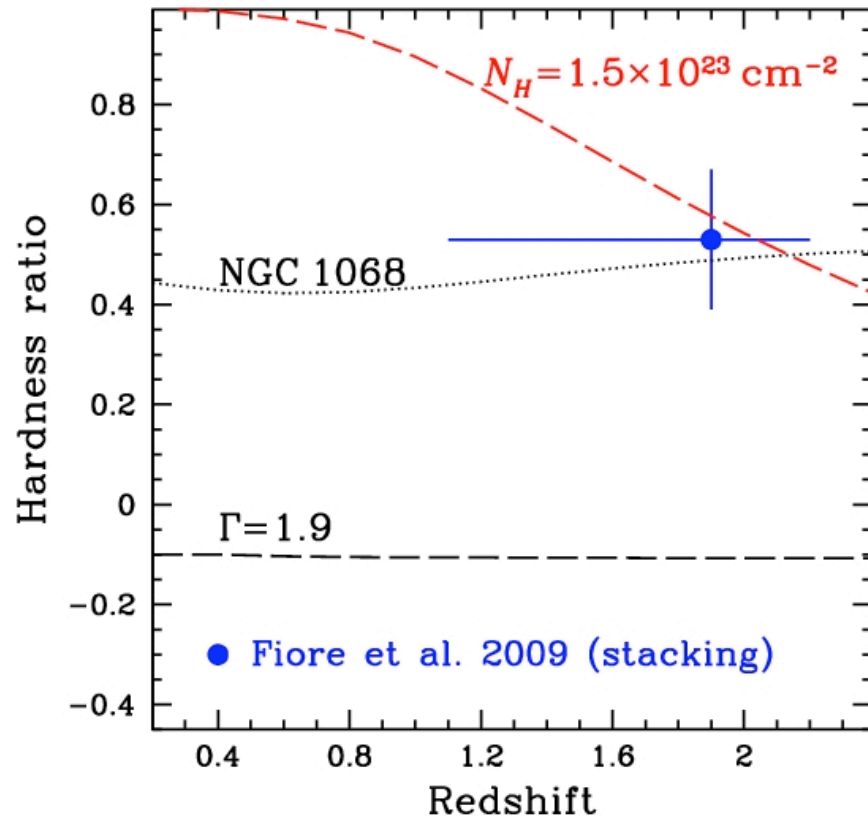


X-ray stacking: hardness ratio consistent with Compton Thick AGN ($N_H > 2 \times 10^{24} \text{cm}^{-2}$).

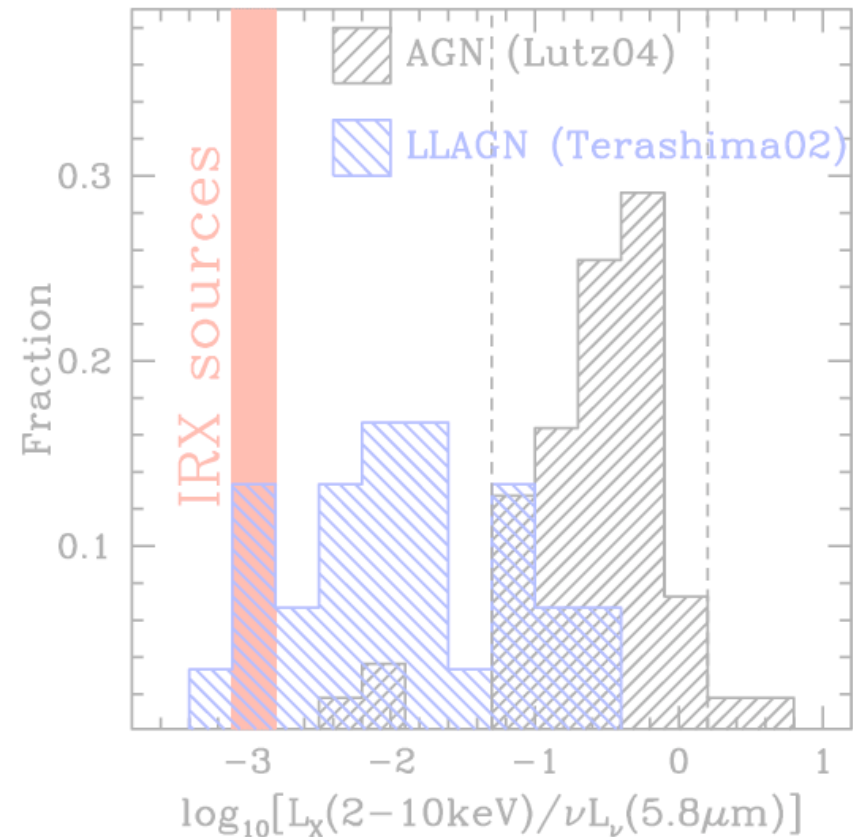


X-ray faint: 2-3dex fainter than typical AGN. Evidence for CT QSOs ($L_X \gg 10^{43} \text{erg/s}$).

IRX sources: Compton Thick QSOs?

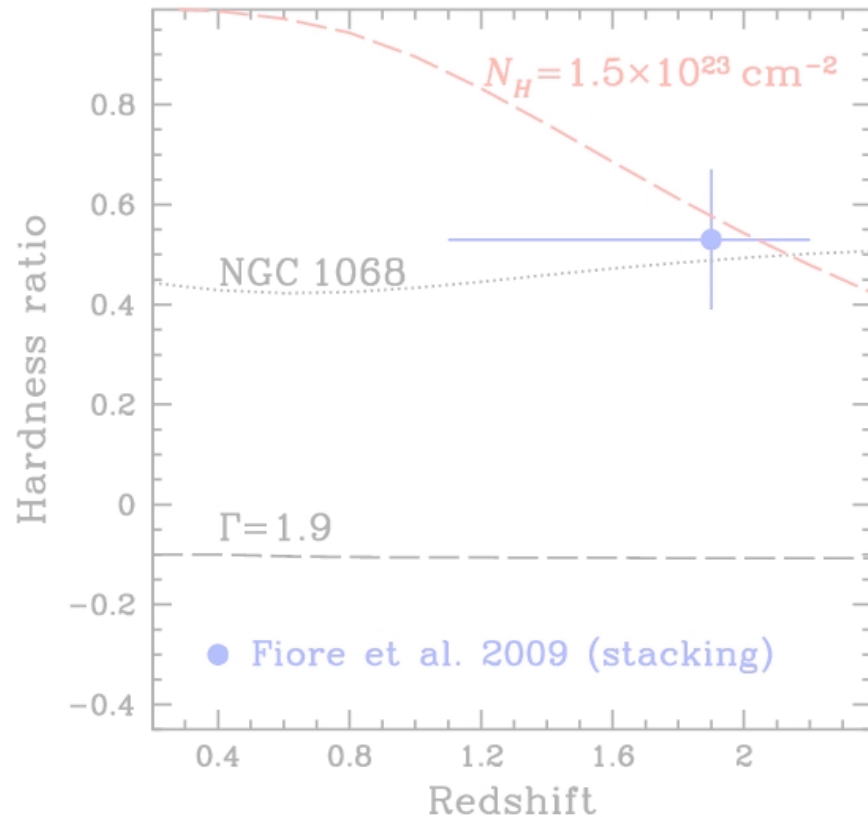


X-ray stacking: hardness ratio also consistent with moderate obscured AGN ($N_H \sim 10^{23} \text{ cm}^{-2}$).

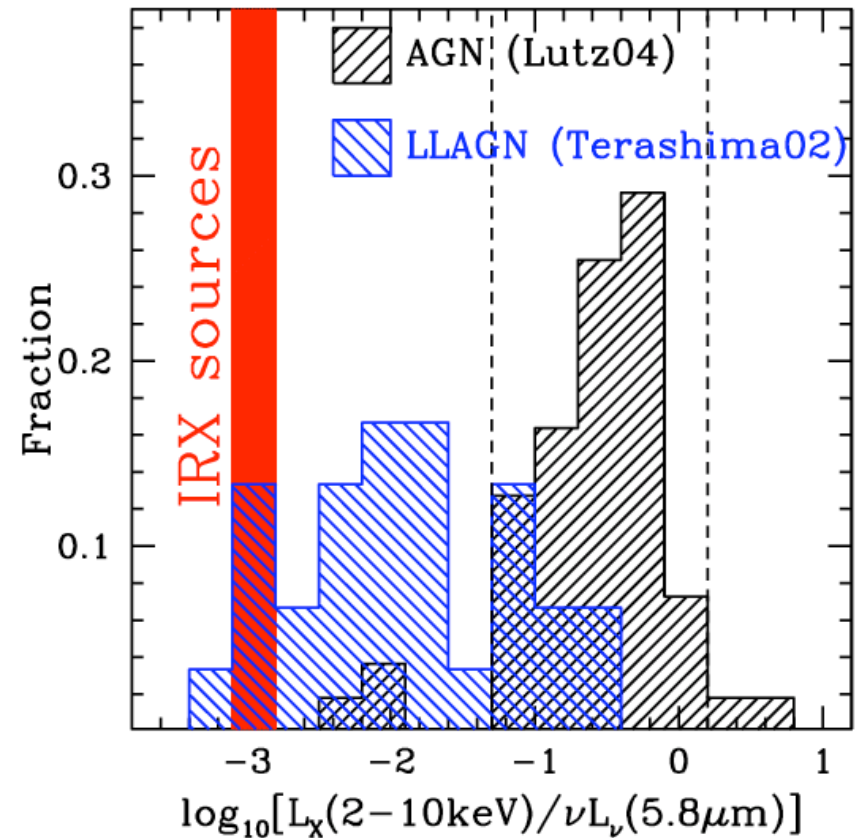


X-ray faint: also consistent with moderate/low luminosity AGN.

IRX sources: Compton Thick QSOs?



X-ray stacking: hardness ratio also consistent with moderate obscured AGN ($N_H \sim 10^{23} \text{ cm}^{-2}$).



X-ray faint: also consistent with moderate/low luminosity AGN.

IRX sources: Compton Thick QSOs?

The nature of IRX sources is **controversial**:

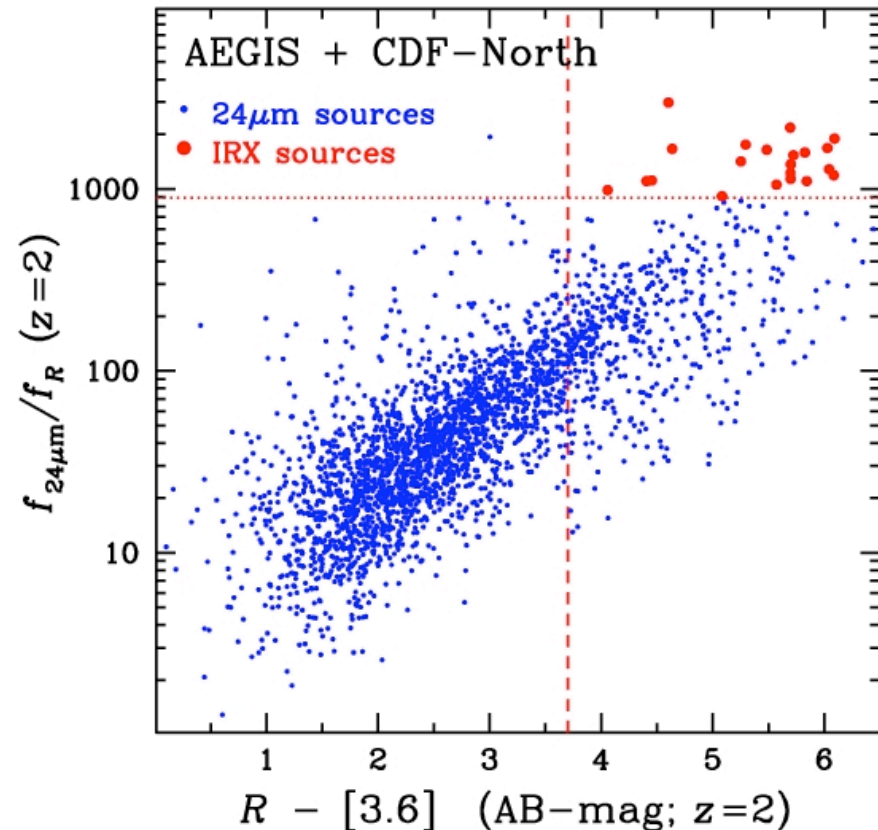
Compton Thick QSOs at $z \sim 2$ (Daddi07, Fiore08, 09)?

OR

Compton Thin moderate- L_X AGN
(Georgantopoulos08, Donley08, Pope09)?

Selection of $z \sim 1$ sources with IRX SEDs

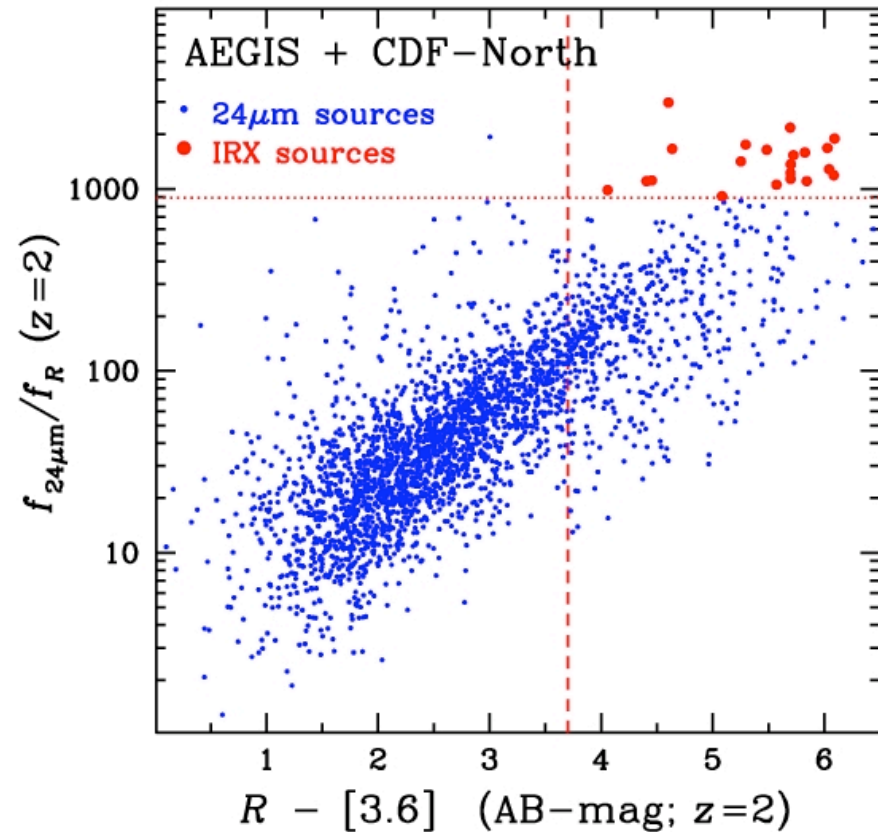
- AEGIS + CDF-North
- UV to far-IR photometry:
RAINBOW* Database
- Select $z \sim 1$ sources with SEDs that satisfy the IRX selection criteria at $z=2$:
 - $f_{24\mu\text{m}}/f_R > 1000$
 - $R - [3.6] > 3.7 \text{ mag}$



* <http://guaix.fis.ucm.es/rainbow>

Selection of $z \sim 1$ sources with IRX SEDs

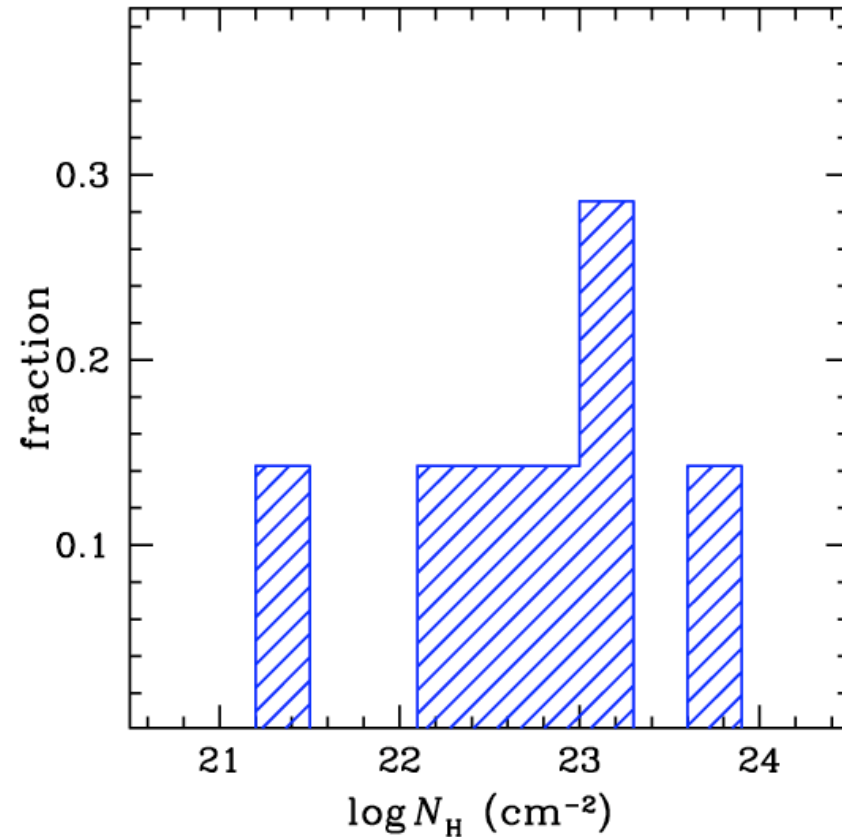
- AEGIS + CDF-North
 - Total of 21 sources with far-UV to $70\mu\text{m}$ SEDs
 - 2-10keV counterparts for 9/21 sources



X-ray detected IRX sources: X-ray spectral analysis

9 IRX sources have
2-10keV counterparts

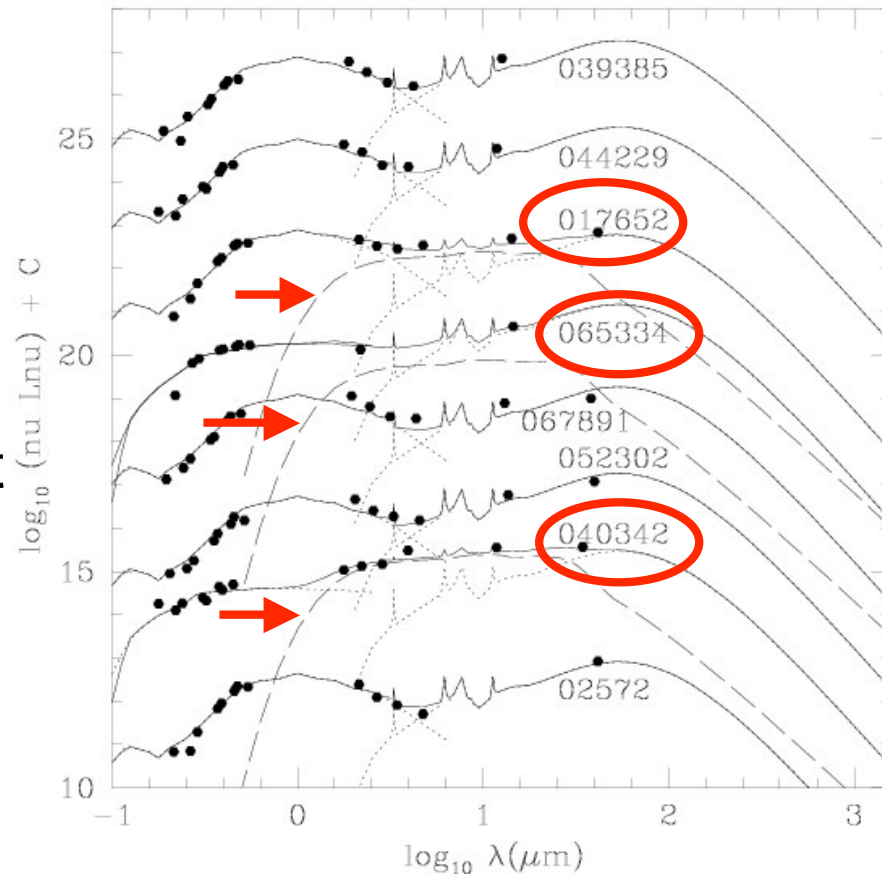
- $N_H \sim 10^{21} - 5 \times 10^{23} \text{ cm}^{-2}$
- $L_X \sim 10^{42} - 2 \times 10^{44} \text{ erg/s}$



X-ray detected IRX sources: SED fits

9 IRX sources have
2-10keV counterparts

- star-formation
contributes to mid-IR
- AGN torus component
is required for *all* X-
ray detected IRXs
with $L_X > 10^{43} \text{ erg/s}$

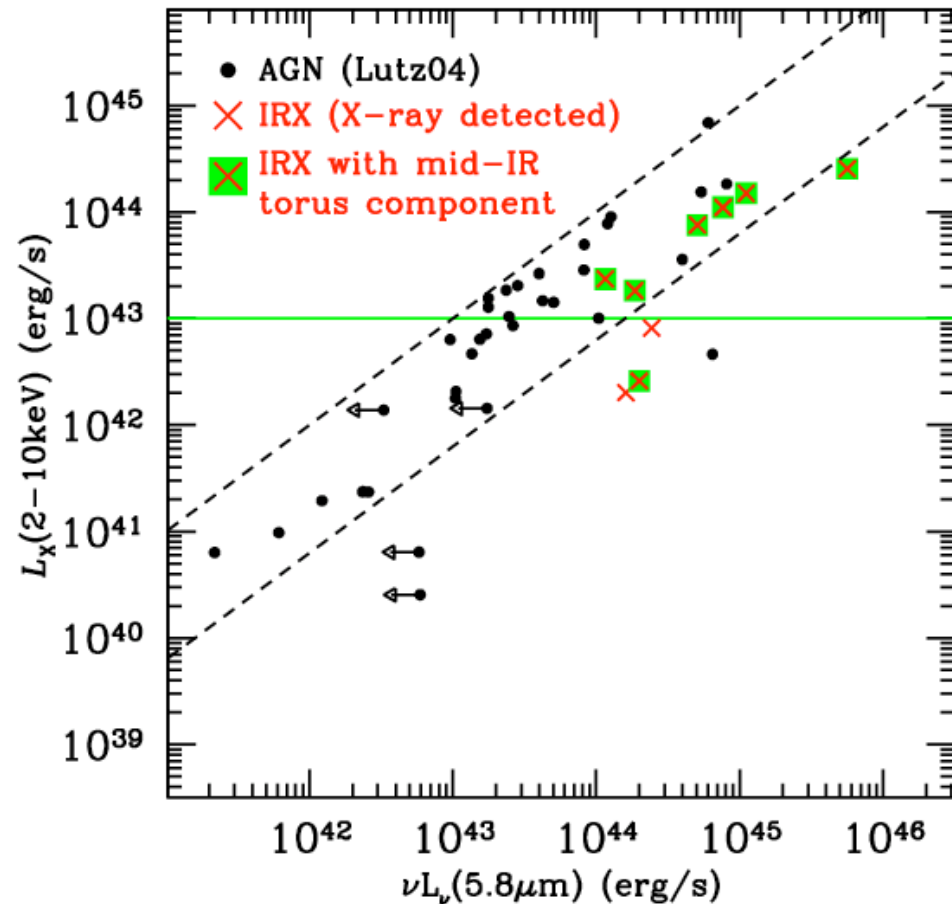


Fit models to SED: Rowan-
Robinson et al. 2005, 2008

X-ray detected IRX sources: SED fits

9 IRX sources have
2-10keV counterparts

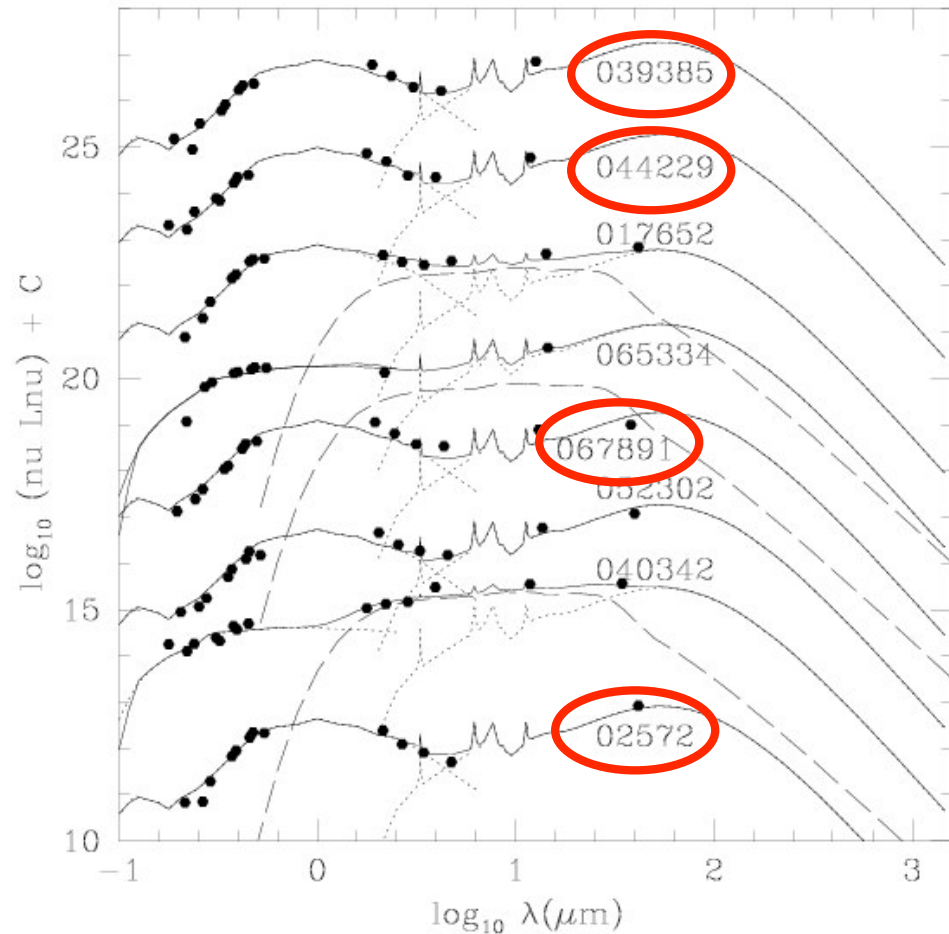
- star-formation
contributes to mid-IR
- QSO torus component
is required for all X-
ray detected IRXs
with $L_X > 10^{43} \text{ erg/s}$



IRX sources w/o X-ray counterparts

12 IRX sources w/o
X-ray detection

- SEDs consistent with starburst activity
- No evidence for AGN component at mid-IR
- AGN intrinsic luminosity $< 10^{43}$ erg/s



Conclusions

- $z \sim 1$ sources with IRX SEDs: no evidence for luminous ($L_X > 10^{43} \text{erg/s}$) Compton Thick ($N_H > 2 \times 10^{24} \text{cm}^{-2}$) AGN.
 - X-ray spectra consistent with moderate column densities ($N_H \sim 10^{23} \text{cm}^{-2}$)
 - Star-formation contributes/dominates the mid-IR.
- The population of IRX sources at $z \sim 2$ is likely to be dominated by Compton thin, moderate/low luminosity AGN.